

1 Random walks, diffusion and polymers

statistical mechanics: bridge between microscopic and macroscopic descriptions

set-up: fluid composed of a large number of identical particles

at any temp $T > 0$: particles subject to thermal motion

we're interested in a test particle (larger than fluid part)

we want to describe the dynamics of this test particle at sufficiently long time scales so that an average large number of collisions with fluid particles has occurred

→ displacement test particle: distance a , randomly

Jump size distribution

random walk = path in space generated by series of steps each selected independently from a given probability distribution $p(\vec{r})$

→ start walk @ origin @ $t=0$, then at times $\Delta t, 2\Delta t, 3\Delta t \dots$ → random steps $\vec{r}_1, \vec{r}_2, \vec{r}_3 \dots$ selected from $p(\vec{r})$

probability distribution $p(\vec{r}) \Rightarrow p(\vec{r}) dx dy dz$ is the prob that $\vec{r} = (x, y, z)$ is selected from interval $[x, x+dx], [y, y+dy], [z, z+dz]$

properties:

- $p(\vec{r}) \geq 0$
- $\langle \vec{r} \rangle = \int p(\vec{r}) \vec{r} d\vec{r} = 0$ (unbiased)
- $\int p(\vec{r}) d\vec{r} = 1$ (normalization)
- $\langle \vec{r}^2 \rangle = \int p(\vec{r}) \vec{r}^2 d\vec{r} = a^2$ (finite value mean squared jump)
- $p(\vec{r})$ has to decay fast @ infinity

examples:

- 1) $p(x) = \frac{1}{2} [\delta(x-a) + \delta(x+a)]$ → RW in a 1D lattice
- 2) $p(x) = \frac{1}{\sqrt{2\pi}a} e^{-x^2/2a^2}$ → RW w/ Gaussian distr. steps
- 3) $p(x) = \begin{cases} \frac{1}{2\sqrt{3}a} & |x| \leq \sqrt{3}a \\ 0 & \text{else} \end{cases}$ → uniform distribution
- 4) $p(\vec{r}) = \frac{1}{4\pi a^2} \delta(|\vec{r}| - a)$ → freely jointed chain (FJC)

End-to-end vector

end-to-end vect. measures the distance of the walk from the origin

→ after N steps: $\langle \vec{R}_N \rangle = \langle \sum_{i=1}^N \vec{r}_i \rangle$
 $= \sum_{i=1}^N \langle \vec{r}_i \rangle = \sum_{i=1}^N \int d\vec{r}_i \vec{r}_i p(\vec{r}_i) = 0$

average displacement is 0 by symmetry (unbiased prop.)

→ squared end-to-end distance:

$\langle \vec{R}_N^2 \rangle = \langle \sum_{i,j=1}^N \vec{r}_i \cdot \vec{r}_j \rangle = \sum_{i=1}^N \langle \vec{r}_i^2 \rangle + \sum_{i \neq j} \langle \vec{r}_i \cdot \vec{r}_j \rangle = \sum_{i=1}^N \langle \vec{r}_i^2 \rangle = a^2 N$
 [since jumps are selected independently and thus $\langle \vec{r}_i \cdot \vec{r}_j \rangle = \langle \vec{r}_i \rangle \cdot \langle \vec{r}_j \rangle = 0$ if $i \neq j$]
 → universal equation (indep. of spec. distr.)

Diffusion equation

end-to-end vector prob. distr.: $P(\vec{R}, t)$ = prob that the walk which started @ origin @ $t=0$ arrives @ $\vec{R}$ @ time t , $t = N\Delta t$

→ $P(\vec{R}, t)$ is a solution of $\frac{\partial P(\vec{R}, t)}{\partial t} = D \nabla^2 P(\vec{R}, t)$ the diffusion equation

$$P(\vec{R}, t + \Delta t) = \int d\vec{r} P(\vec{R} - \vec{r}, t) p(\vec{r}) \quad [\text{RW in a d-dim. space}]$$

= prob of being in $\vec{R} - \vec{r}$ @ t multiplied by the prob of making a jump of $\vec{r}$
 → prod of probs since each new jump is indep from the previous ones

$$\approx \int d\vec{r} \left\{ P(\vec{R}, t) - \vec{r} \cdot \nabla P(\vec{R}, t) + \frac{1}{2} \sum_{k,l} r_k r_l \frac{\partial^2 P(\vec{R}, t)}{\partial R_k \partial R_l} \right\} p(\vec{r})$$

(expansion in $\vec{r}$ & considered lower order terms)

$$= P(\vec{R}, t) + \frac{\sigma^2}{2d} \nabla^2 P(\vec{R}, t)$$

(und. isotropy property for d-dim walks: $\langle r_k r_l \rangle = \delta_{kl} \frac{\sigma^2}{d}$)

→ for small Δt :

$$\frac{\partial P(\vec{R}, t)}{\partial t} = \frac{\sigma^2}{2d \Delta t} \nabla^2 P(\vec{R}, t)$$

→ diffusion eq. with $D = \frac{\sigma^2}{2d \Delta t}$

↳ describes the behaviour of the collective motion of the fluid particles resulting from the random walk movements of each particle → diffusion of particles

↳ the presumably random moving of particles suspended in a fluid resulting from their bombardment by the fast-moving particles (atmol.) in the fluid

→ M non-interacting, Brownian part concentration $c(\vec{R}, t) = M P(\vec{R}, t)$

• a RW describing the traj. of a part in time

→ zig-zag motion = Brownian motion

• the diffusion eq. describes the time evolution of the concentration of particles spreading on some medium, each indiv. part performing a Brownian motion

• rewrite diff. eq: $\frac{\partial c(\vec{R}, t)}{\partial t} = -\vec{\nabla} \cdot \vec{j}(\vec{R}, t)$

with current $\vec{j}(\vec{R}, t) = -D \nabla c(\vec{R}, t)$

→ continuity eq.: describes time evolution conserved quantity, here: $\int d\vec{R} P(\vec{R}, t)$ is its in time (conservation of particles)

• reaction-diff equation

$$\frac{\partial c(\vec{R}, t)}{\partial t} = D \nabla^2 c(\vec{R}, t) + f(c(\vec{R}, t))$$

describes chemical reactions

qual → total concentration not conserved
 → $f(c)$ a non-linear function

• quinal form diff eq

$$\frac{\partial c(\vec{R}, t)}{\partial t} = \vec{\nabla} \cdot \{ D(c(\vec{R}, t)) \vec{\nabla} c(\vec{R}, t) \}$$

→ diff coeff D can depend on concentration and position

Solving the diffusion equation

• Gaussians

$$G_{\vec{R}_0}(\vec{R}, t) = \left(\frac{1}{4\pi Dt} \right)^{d/2} e^{-\frac{(\vec{R} - \vec{R}_0)^2}{4Dt}}$$

$$\lim_{t \rightarrow 0} G_{\vec{R}_0}(\vec{R}, t) = \delta(\vec{R} - \vec{R}_0)$$

$$\text{and } \sigma^2 = 2Dt \quad (\sigma = \sqrt{2Dt})$$

start from localized δ at $t=0$ (peak)
 spread out later in time $\sim \sqrt{t}$

↳ linear combinations $G_{\vec{R}_0}$ also solutions since diff eq is linear

$$\text{most general solution: } c(x, t) = \int dy c(y) G_y(x, t)$$

$$\text{with right BC: } \lim_{t \rightarrow 0} c(x, t) = c(x) = \int dy c(y) \delta(y - x)$$

$$g(x) = \frac{1}{\sigma \sqrt{2\pi}} e^{-\frac{1}{2} \left(\frac{x - \mu}{\sigma} \right)^2} \quad (\text{for 1D})$$



Drift-diffusion equation

• external force acting on diff part → additional current j_F

$$j_F = v_d(x) c(x, t) = \frac{1}{\gamma} c(x, t) F(x) = -\frac{1}{\gamma} c(x, t) \frac{dV}{dx}$$

with v_d = drift velocity, γ = friction coeff, $V(x)$ the potential, since $F = -\gamma v$

→ drift-diff eq: $\frac{\partial c(x, t)}{\partial t} = -\frac{\partial j_{\text{tot}}}{\partial x} = D \frac{\partial^2 c(x, t)}{\partial x^2} + \frac{1}{\gamma} \frac{\partial}{\partial x} \left\{ c(x, t) \frac{dV}{dx} \right\}$

friction/viscous forces dominate
 neglect initial acceleration

• in total equilibrium: net total current vanishes

$$j_{\text{tot}} = j_D + j_F = -D \frac{dc_{\text{eq}}(x)}{dx} - \frac{1}{\gamma} c_{\text{eq}}(x) \frac{dV}{dx} = 0$$

$$\text{with } c_{\text{eq}}(x) = A \exp\left(-\frac{V(x)}{k_B T}\right) \quad \text{Journ solution}$$

T temp, k_B Boltzmann, A norm.

↳ distr expected for non-interacting part in eq. @ temp T

$$\Rightarrow D = \frac{k_B T}{\gamma}$$

Einstein relation

drift-diff

$$\frac{\partial c(x, t)}{\partial t} = D \frac{\partial}{\partial x} \left\{ e^{-\beta V(x)} \frac{\partial}{\partial x} (e^{\beta V(x)} c(x, t)) \right\}$$

$$D \frac{\partial}{\partial x} \left(e^{-\beta V(x)} \left(e^{\beta V(x)} \frac{\partial c(x, t)}{\partial x} + e^{\beta V(x)} c(x, t) \frac{\partial V}{\partial x} \right) \right)$$

2) Ensembles in Classical Statistical Mechanics

system of N particles for which position and momenta are given (microscopic)

Introduction

microscopic state of a syst. defined by a $6N$ -dim vector $\Gamma \equiv (\vec{p}_1, \vec{p}_2, \dots, \vec{p}_N, \vec{q}_1, \vec{q}_2, \dots, \vec{q}_N)$

$\rightarrow \Gamma$ -space = $6N$ dim vector space

Hamiltonian $\mathcal{H}$ governs time-evolution

$$\mathcal{H}(\Gamma) = \sum_{i=1}^N \frac{\vec{p}_i^2}{2m} + \Phi(\vec{q}_1, \vec{q}_2, \dots, \vec{q}_N)$$

observables $A(\Gamma)$ (functions in Γ -space)

ex. kin. energy $A(\Gamma) = \sum_{i=1}^N \vec{p}_i^2 / 2m$

tot. energy $A(\Gamma) = \mathcal{H}(\Gamma)$

part. density $A(\Gamma, \vec{r}) = \sum_{i=1}^N \delta(\vec{q}_i - \vec{r})$

part @ a given pos.; for a homogeneous syst

$\langle \rho(\vec{r}, t) \rangle$ a constant indep. on $\vec{r}$

$\rightarrow$ average values

* pure $\langle A \rangle_{\text{pure}} \equiv \frac{1}{\tau} \int_{t_0}^{t_0+\tau} A(\Gamma(t)) dt$

(time average; with $\Gamma(t)$ a path in Γ -space, solution to eq. of motion)

$\sim$ experiments:

$\rightarrow$ if equil.: $\langle A \rangle$ indep on t_0 and τ

t_0 = time @ which measurement started, τ = duration $\rightarrow \langle A \rangle$ = average measured value observable

* SM $\langle A \rangle = \int d\Gamma \rho(\Gamma) A(\Gamma)$

ensemble average, using prob. distr. func. $\rho(\Gamma)$ which represents the prob. of finding a system in a microstate characterized by Γ

3 different types of $\rho(\Gamma) \rightsquigarrow$ microcanonical ensemble

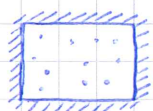
canonical ensemble

grandcanonical ensemble

in the limit of $N \rightarrow \infty \rightarrow$ all $\rho(\Gamma)$ lead to the same value of thermal averages

$\hookrightarrow$ in principle we can use any of the 3 ensembles to calc. thermal av., in practice canonical more convenient

Microcanonical ensemble



E, V, N

isolated syst with fixed volume V , fixed #part N and no energy exchange

$\rightarrow$ total mechanical energy E conserved

$\rightarrow$ part trajectories found in manifolds of constant tot. energy, ex. (MO)

for a fixed energy, time is constant along a part path in Γ

* PRINCIPLE OF EQUAL A PRIORI PROBABILITY (SM, general)

All the microstates with a constant E are equally probable

$\rightarrow$ prob. dens $\rho_{mc}(\Gamma) = \frac{\delta(E - \mathcal{H}(\Gamma))}{\omega(E, N, V) N! h^{3N}}$

N = # part, h = Planck's constant

$$= \frac{1}{N! h^{3N}} \frac{\delta(E - \mathcal{H}(\Gamma))}{\int \frac{d\Gamma}{N! h^{3N}} \delta(E - \mathcal{H}(\Gamma))}$$

a.p. principle of indistinguishability of identical particles

$\rightarrow$ necessary for extremum thermodyn. quantities

$$\omega(E) = \int \frac{d\Gamma}{N! h^{3N}} \delta(E - \mathcal{H}(\Gamma))$$

microcanonical dens. of states

$\sim 6N-1$ dim "volume" of manifold $E = \text{cte}$ in Γ -space

$$\rightarrow \Omega(E, N, V) = \int \frac{d\Gamma}{N! h^{3N}} \theta(E - \mathcal{H}(\Gamma))$$

$\sim$ volume of region $\mathcal{H}(\Gamma) \leq E$



$\theta(x) = \begin{cases} 1 & x \geq 0 \\ 0 & x < 0 \end{cases}$ step function

with $\omega(E, N, V) = \frac{\partial \Omega(E, N, V)}{\partial E}$

$$\approx \frac{\Omega(E) - \Omega(E - \Delta E)}{\Delta E}$$

for $N \gg 1$, $N \sim 10^{23}$ nearly else

- connecting microscopic to macroscopic

BOLTZMANN $S(E, N, V) = k_B \log \omega(E, N, V)$

entropy $\approx k_B \log \Omega(E, N, V)$

$$\rightarrow \Omega(E, N, V) \approx \omega(E, N, V) \Delta E$$

example: Ideal gas

$$\mathcal{H}(\Gamma) = \sum_i \frac{\vec{p}_i^2}{2m} + \sum_i \varphi_{\text{wall}}(\vec{q}_i)$$

$$\text{with } \varphi_{\text{wall}}(\vec{q}_i) = \begin{cases} 0 & \text{if } \vec{q}_i \in V \\ +\infty & \text{else} \end{cases}$$

→ confinement (fixed volume microcan.)

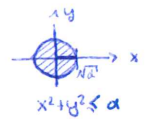
$$\rightarrow \Omega(E, N, V) = \int \frac{d\Gamma}{N! h^{3N}} \Theta(E - \sum_i \frac{\vec{p}_i^2}{2m} - \sum_i \varphi_{\text{wall}}(\vec{q}_i))$$

$$= \frac{1}{N! h^{3N}} \int_V d\vec{q}_1 \int_V d\vec{q}_2 \dots \int_V d\vec{q}_N \int d\vec{p}_1 \dots d\vec{p}_N \Theta(E - \sum_i \frac{\vec{p}_i^2}{2m})$$

$$= \frac{1}{N! h^{3N}} V^N \int d\vec{p}_1 \dots d\vec{p}_N \Theta(E - \sum_i \frac{\vec{p}_i^2}{2m})$$

$$= \frac{V^N}{N! h^{3N}} \frac{\pi^{3N/2}}{(\frac{3N}{2})!} (2mE)^{3N/2}$$

$$\int dx dy \Theta(a - (x^2 + y^2)) = \pi a$$



for d-dim

$$V_d(R) = \frac{\pi^{d/2}}{(d/2)!} R^d \quad \text{and } \frac{1}{2}! = \frac{\sqrt{\pi}}{2} \rightarrow \frac{3}{2}! = \frac{3\sqrt{\pi}}{2}$$

3N-dim. with $R = \sqrt{2mE}$

(Boltzmann) $\rightarrow S \approx k_B \log \Omega(E, N, V)$

$$P = T \frac{\partial S}{\partial V} \Big|_E = T \frac{\partial}{\partial V} k_B (\log V^N + \dots) \Big|_E$$

$$= T k_B \frac{N}{V} \rightarrow \text{ideal gas law}$$

$$S \approx k_B \log \left[\frac{V^N}{N! h^{3N}} \left(\frac{2\pi m E}{3N} \right)^{3N/2} \frac{1}{E^{-3N/2}} \right]$$

using Stirling: $N! \approx N^N e^{-N}$

$$\left(\frac{3N}{2}\right)! \approx \left(\frac{3N}{2}\right)^{3N/2} e^{-3N/2}$$

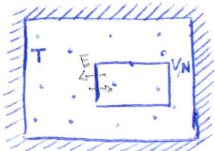
$$= N k_B \log \left[\frac{V}{N} \left(\frac{4\pi m E}{3N h^2} \right)^{3/2} \right] + \frac{5N}{2} k_B$$

→ extensivity shows that $S(\alpha E, \alpha V, \alpha N) = \alpha S(E, V, N)$

output extensivity $\rightarrow N! \rightarrow$ all origin (indistinguishability particles)

intensive	extensive
P	V applied P adjusts V
μ	N
T	S
...	E ...

canonical ensemble



system in contact with a thermal bath fixed at temp T

N, V also fixed

system and bath can exchange energy $E \rightarrow$ no longer fixed

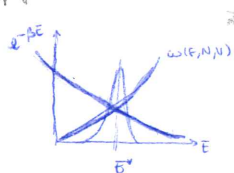
$$\rightarrow \text{prob. dens: } \rho_c(\Gamma) = \frac{e^{-\beta \mathcal{H}(\Gamma)}}{N! h^{3N} \int \frac{d\Gamma}{N! h^{3N}} e^{-\beta \mathcal{H}(\Gamma)}} = \frac{e^{-\beta \mathcal{H}(\Gamma)}}{N! h^{3N} Z(N, V, T)}$$

$$\beta = \frac{1}{k_B T}$$

$$F(\omega) = \int_0^\infty f(t) e^{-\omega t} dt \quad \text{with } Z(N, V, T) = \int \frac{d\Gamma}{N! h^{3N}} e^{-\beta \mathcal{H}(\Gamma)}$$

Laplace transfo of $f(t)$

Laplace transfo $\rightarrow$



$$Z(\beta) = \int_0^\infty e^{-\beta E} \omega(E, N, V) dE$$

now for $N, V, T \rightarrow \infty$ (TD limit)

$\omega(E, N, V)$ part growing function E

$e^{-\beta E}$ part decaying function E

$\rightarrow$ peaked around char value E^*

$$\Rightarrow Z(N, V, T) \approx e^{-\beta E^*} \omega(E^*, N, V)$$

$$= e^{-\beta E^*} e^{S(E^*, N, V)/k_B} \quad \text{Boltzmann assumption}$$

$$= e^{-\beta [E^* - TS(E^*, N, V)]} = e^{-\beta F(E^*, N, V)}$$

or use saddle point approx (see next page)

$$\Rightarrow F = E - TS \approx -k_B T \log Z \quad \text{(Helmholtz free energy)}$$

now for non-interacting particles

$$\mathcal{H}(\Gamma) = \sum_{i=1}^N \mathcal{H}_1(\Gamma_i) = \sum_{i=1}^N \left(\frac{\vec{p}_i^2}{2m} + \varphi(\vec{q}_i) \right)$$

$$\Rightarrow Z(N, V, T) = \frac{Z_1(N, V, T)^N}{N!}$$

$$\text{with } Z_1 = \int d\vec{p} d\vec{q} e^{-\beta \mathcal{H}_1(\vec{p}, \vec{q})}$$

and in general, since the momenta integrals are gaussian integrals

$$Z(N, V, T) = \frac{1}{N!} \lambda_T^{-3N} \int d\vec{q}_1 \dots d\vec{q}_N e^{-\beta \varphi(\vec{q}_1, \dots, \vec{q}_N)} \equiv \frac{Q(N, V, T)}{N! \lambda_T^{3N}}$$

with $Q(N, V, T)$ the configurational partition function

$$\text{and } \lambda_T = \frac{h}{\sqrt{2\pi m k_B T}} \quad \text{the thermal wavelength}$$

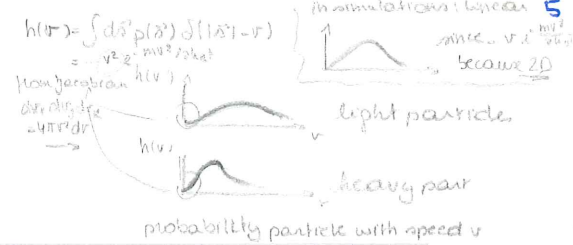
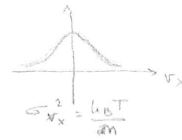
Gaussian integr.

$$\int_{-\infty}^{+\infty} e^{-a(x+b)^2} dx$$

$$= \sqrt{\frac{\pi}{a}}$$

$$\int_{-\infty}^{+\infty} e^{-\frac{1}{2} \frac{p^2}{m k_B T}} dp$$

$$= \sqrt{2\pi m k_B T}$$



momenta gaussian diste $\Rightarrow$ averages $\langle A(\vec{p}) \rangle = \langle A(\vec{v}) \rangle$ trivial
- diste. velocities follows a universal law indep. on type of interaction

* MAXWELL DISTRIBUTION OF VELOCITIES

$$p(\vec{v}) = \left(\frac{m}{2\pi k_B T} \right)^{3/2} e^{-m\vec{v}^2 / 2k_B T}$$

(prob diste.)

$$\Rightarrow E = \langle \mathcal{H}(\Gamma) \rangle = - \frac{\partial \log Z(N, V, T)}{\partial \beta}$$

internal energy

or average energy

$$\begin{aligned} \langle A \rangle &= \frac{\int d\Gamma p(\Gamma) A(\Gamma)}{\int d\Gamma p(\Gamma)} \\ &= \frac{\int d\Gamma e^{-\beta \mathcal{H}(\Gamma)} A(\Gamma)}{\int d\Gamma e^{-\beta \mathcal{H}(\Gamma)}} \end{aligned}$$

with $p(\Gamma)$ norm.: $\int d\Gamma p(\Gamma) = 1$

$$\Rightarrow \langle \mathcal{H} \rangle = \frac{\int d\Gamma e^{-\beta \mathcal{H}(\Gamma)} \mathcal{H}(\Gamma)}{\int d\Gamma e^{-\beta \mathcal{H}(\Gamma)}} = - \frac{\partial}{\partial \beta} \log \left(\int d\Gamma e^{-\beta \mathcal{H}(\Gamma)} \right)$$

$$\text{then } \frac{\partial}{\partial \beta} \log \left(\int d\Gamma e^{-\beta \mathcal{H}(\Gamma)} \right) = \frac{1}{\int d\Gamma e^{-\beta \mathcal{H}(\Gamma)}} \int d\Gamma (-\mathcal{H}) e^{-\beta \mathcal{H}(\Gamma)}$$

$$\text{and } Z(N, V, T) = \int \frac{d\Gamma}{N! h^{3N}} e^{-\beta \mathcal{H}(\Gamma)}$$

$$\text{thus } \langle \mathcal{H} \rangle = - \frac{\partial}{\partial \beta} \log Z$$

• saddle point approx

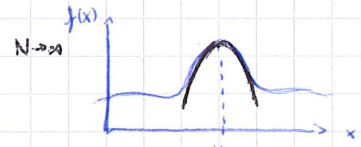
to go from $Z(N, V, T) = \int dE e^{-\beta E} \omega(E, N, V)$ to $Z(N, V, T) \approx e^{-\beta E^*} \omega(E^*, N, V)$

approx as follows

$$I = \int dx e^{N f(x)} \approx e^{N f(x_0)} \int_{-\infty}^{\infty} dx e^{\frac{N}{2} f''(x_0) (x-x_0)^2}$$

$$\xrightarrow{N \rightarrow \infty} e^{N f(x_0)} \sqrt{\frac{2\pi}{N |f''(x_0)|}}$$

dominant part



dominant part is the maximum

$$\Rightarrow \log I = N f(x_0) + O(\log N)$$

$$\Rightarrow \text{taking } \log Z = \beta(E - TS) \text{ is a good approx.}$$

• thermal wavelength

$$\lambda_T = \frac{h}{\sqrt{2\pi m k_B T}} = \frac{h}{p_T}$$

~ average de Broglie wavelength gas part in ideal gas at specified T
 $\rightarrow$ the Broglie wavelength for a part. with average thermal energy

comparing the wavelength λ_T with the typical distance betw. part $\sim \frac{1}{N}$

$\Rightarrow$ if $\lambda_T \ll \frac{1}{N} \rightarrow$ classical regime

if $\lambda_T \gtrsim \frac{1}{N} \rightarrow$ quantum regime

$$\frac{\lambda_T}{\frac{1}{N}} \ll 1 \quad \frac{\lambda_T}{\frac{1}{N}} \gtrsim 1$$

• non-interacting particles

$$Z(N, V, T) = \frac{1}{N! \lambda_T^{3N}} \int d\vec{q}_1 \dots d\vec{q}_N e^{-\beta \varphi(\vec{q}_1)} e^{-\beta \varphi(\vec{q}_2)} \dots e^{-\beta \varphi(\vec{q}_N)}$$

$$= \frac{1}{N! \lambda_T^{3N}} \int d\vec{q}_1 e^{-\beta \varphi(\vec{q}_1)} \int d\vec{q}_2 e^{-\beta \varphi(\vec{q}_2)} \dots$$

$$= \frac{1}{N! \lambda_T^{3N}} \left[\int d\vec{q} e^{-\beta \varphi(\vec{q})} \right]^N = \frac{[Q(1, V, T)]^N}{N! \lambda_T^{3N}}$$

$\rightarrow$ in a gravitational field

$$Q(1, V, T) = \int d\vec{q}_1 \dots d\vec{q}_N e^{-\beta m g (z_1 + z_2 + \dots + z_N)}$$

$$= \int dx_1 dy_1 dz_1 e^{-\beta m g z_1} \int dx_2 dy_2 dz_2 e^{-\beta m g z_2} \dots$$

$$= \left[\int_0^L dx \int_0^L dy \int_0^L dz e^{-\beta m g z} \right]^N = \left[L^2 \frac{1}{\beta m g} (1 - e^{-\beta m g L}) \right]^N$$



example: Ideal gas

$$Z = \frac{1}{N! h^{3N}} \int d\vec{r} e^{-\beta \left[\sum_i \frac{p_i^2}{2m} + \sum_i \varphi(\vec{r}_i) \right]} = \frac{1}{N! h^{3N}} \left[\int d\vec{p} d\vec{r} e^{-\beta \left(\frac{p^2}{2m} + \varphi(\vec{r}) \right)} \right]^N$$

same φ as before

$$= \frac{V^N}{N! h^{3N}} \left[\int d\vec{p} e^{-\beta p^2 / 2m} \right]^N = \frac{V^N}{N! h^{3N}} \left[\int dp e^{-\beta p^2 / 2m} \right]^{3N} = \frac{V^N}{N! h^{3N}} (2\pi m / \beta)^{3N/2}$$

$$= \frac{V^N}{N! \lambda_T^{3N}}$$

$$\chi = \frac{V^N}{N! \lambda_T^{3N}}$$

$$\rightarrow F = -k_B T \log \chi = -k_B T [N \log V - 3N \log \lambda_T - \log N!]$$

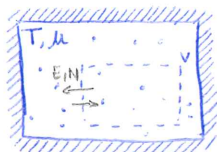
$$P = \left. \frac{\partial F}{\partial V} \right|_{N,T} = \frac{N k_B T}{V} \rightarrow \text{ideal gas law}$$

$$\lambda_T = \frac{h}{\sqrt{2\pi m k_B T}} = A \beta^{1/2}$$

$$\text{and } E = \langle H \rangle = - \frac{\partial \log \chi}{\partial \beta} = \frac{\partial (3N \log \beta^{1/2})}{\partial \beta} = \frac{3N}{2} \frac{\partial \log \beta}{\partial \beta} = \frac{3}{2} N k_B T \quad \text{OK!}$$

= average kinetic energy of a system

Grand canonical ensemble



V, T, μ (chemical potential) fixed
 E, N can fluctuate

→ prob dens

$$\rho(\Gamma_N, N) = \frac{e^{-\beta \mathcal{H}(\Gamma_N)} e^{\beta \mu N}}{N! h^{3N} \Xi(\mu, V, T)}$$

$$\sum_N \int d\Gamma_N \rho_{GC}(\Gamma_N, N) = 1$$

$$\text{with } \Xi(\mu, V, T) = \sum_N e^{\beta \mu N} \int \frac{d\Gamma_N}{N! h^{3N}} e^{-\beta \mathcal{H}(\Gamma_N)}$$

Laplace transform

$$= \sum_N e^{\beta \mu N} \chi(N, V, T) = \sum_N e^{\beta \mu N} \int dE e^{-\beta E} \omega(E, N, V)$$

sum dominated by N_* the most probable value of the # of particles

$$\approx e^{\beta \mu N_*} \chi(N_*, V, T)$$

$$= e^{\beta \mu N_*} e^{-\beta F} = e^{-\beta [F - \mu N_* + E - TS]} = e^{\beta PV}$$

$$\Rightarrow k_B T \log \Xi = PV$$

we can also show that

$$E = \langle \mathcal{H}(\Gamma) \rangle = - \left. \frac{\partial \log \Xi(N, V, T)}{\partial \beta} \right|_{\beta \mu = \text{const}}$$

$$\langle N \rangle = \frac{\partial \log \Xi(N, V, T)}{\partial \mu}$$

$$\langle \mathcal{H}(\Gamma) \rangle = \frac{\sum_N \mathcal{H}(\Gamma_N) e^{\beta \mu N} \chi(N, V, T)}{\sum_N e^{\beta \mu N} \chi(N, V, T)} = - \frac{\partial \log \chi}{\partial \beta} = - \left. \frac{\partial \log \Xi}{\partial \beta} \right|_{\beta \mu = \text{const}}$$

$$\langle N \rangle = \frac{\sum_N N e^{\beta \mu N} \chi(N, V, T)}{\sum_N e^{\beta \mu N} \chi(N, V, T)} = \frac{1}{\beta} \frac{\partial \log \Xi}{\partial \mu} = \frac{1}{\beta} \frac{\partial \Xi}{\partial \mu} \quad \text{OK!}$$

where

$$E = TS - PV + \mu N$$

$$\rightarrow \mu N = E - TS + PV = G$$

(Euler relation)

example: Ideal gas

$$\Xi = \sum_N e^{\beta \mu N} \chi(N, V, T) = \sum_N e^{\beta \mu N} \frac{V^N}{N! \lambda_T^{3N}} = \sum_N \left(\frac{e^{\beta \mu V}}{\lambda_T^3} \right)^N \frac{1}{N!}$$

$$= \exp \left(\frac{e^{\beta \mu V}}{\lambda_T^3} \right) \quad \lambda_T \propto \beta^{1/2}$$

$$\exp(x) = \sum_{k=0}^{\infty} \frac{x^k}{k!}$$

$$\Rightarrow \frac{PV}{k_B T} = \log \Xi = \frac{e^{\beta \mu V}}{\lambda_T^3} \rightarrow \text{how does this relate to the ideal gas law } PV = N k_B T?$$

$$\langle N \rangle = \frac{\partial \log \Xi}{\partial \mu} = \frac{1}{\beta} \frac{\partial}{\partial \mu} \left(\frac{e^{\beta \mu V}}{\lambda_T^3} \right) = \frac{e^{\beta \mu V}}{\lambda_T^3}$$

$$\rightarrow \frac{PV}{k_B T} = \langle N \rangle \quad \text{OK!}$$

$$E = - \left. \frac{\partial \log \Xi}{\partial \beta} \right|_{\beta \mu = \text{const}} = - \frac{\partial}{\partial \beta} \left(\frac{e^{\beta \mu V}}{\lambda_T^3} \right) \Big|_{\beta \mu = \text{const}} = - e^{\beta \mu V} \frac{\partial}{\partial \beta} \left(\frac{1}{\lambda_T^3} \right) = \frac{3}{2} e^{\beta \mu V} V \sqrt{\frac{2\pi m}{h^2}} \sqrt{\beta}^{-3/2}$$

$$= \frac{3}{2} \frac{1}{\beta} \frac{e^{\beta \mu V}}{\lambda_T^3} = \frac{3}{2} k_B T \langle N \rangle$$

→ Ideal gas

next with non-interacting particles: $\mathcal{H} = \sum_{i=1}^N \frac{\vec{p}_i^2}{2m} + \sum_{i=1}^N \psi_w(\vec{q}_i)$

micro-canonical

$$\omega(E, N, V) = \frac{V^N}{N! h^{3N}} \frac{(2\pi m E)^{3N/2}}{(3N/2 - 1)!}$$

$$\sim \left(\frac{V}{N} \right)^N \left(\frac{4\pi m E}{3h^2 N} \right)^{3N/2} e^{5N/2}$$

$$\rightarrow p_{mc}(\vec{p}) = \langle \delta(\vec{p} - \vec{p}_1) \rangle_{mc}$$

$$\sim \dots \exp \left(- \frac{3N \vec{p}^2}{4mE} \right) \quad \text{TODO} \rightarrow p_{41}$$

canonical

$$\rightarrow Z(N, V, T) = \frac{V^N}{N! \lambda_T^{3N}}$$

$$\rightarrow p_c(\vec{p}) = \langle \delta(\vec{p} - \vec{p}_1) \rangle_c = \dots e^{-\beta \frac{\vec{p}^2}{2m}} \quad \text{TODO} \rightarrow p_{42}$$

⇒ you'll see that with the 3 ensembles, the ideal gas law $PV = N k_B T$ and the internal energy $E = \frac{3}{2} N k_B T$ can be calculated (see throughout summary)

if $N \rightarrow \infty$:
mc, c and gc
all become the same

Equipartition theorem

* $\left\langle x_i \frac{\partial \mathcal{H}}{\partial x_i} \right\rangle = k_B T \delta_{i,3}$ with x_i a comp. of pos. or mom. of the particle
↳ thermal average

$\rightarrow \langle q_{i,x} \frac{\partial \mathcal{H}}{\partial q_{i,y}} \rangle = \langle q_{i,x} \frac{\partial \mathcal{H}}{\partial q_{i,x}} \rangle = \langle p_{i,x} \frac{\partial \mathcal{H}}{\partial p_{i,x}} \rangle = 0$ for $i \neq j$

now

$\langle \vec{p}_i \cdot \frac{\partial \mathcal{H}}{\partial \vec{p}_i} \rangle = \langle p_{i,x} \frac{\partial \mathcal{H}}{\partial p_{i,x}} \rangle + \langle p_{i,y} \frac{\partial \mathcal{H}}{\partial p_{i,y}} \rangle + \langle p_{i,z} \frac{\partial \mathcal{H}}{\partial p_{i,z}} \rangle = 3k_B T$

$\Rightarrow$ total kinetic energy (summing over all N part): $E_{kin} = \frac{3Nk_B T}{2}$

since $\langle \vec{p} \cdot \frac{\partial \mathcal{H}}{\partial \vec{p}} \rangle = \langle \vec{p} \cdot \frac{\vec{p}}{m} \rangle = 2 \langle \frac{p^2}{2m} \rangle = 2 \langle \mathcal{H} \rangle$

↳ since $E_{kin} = E_{ideal\ gas}$ we can conclude that the only contribution comes from the kinetic energy

and

$\langle \vec{q}_i \cdot \frac{\partial \mathcal{H}}{\partial \vec{q}_i} \rangle = \langle q_{i,x} \frac{\partial \mathcal{H}}{\partial q_{i,x}} \rangle + \langle q_{i,y} \frac{\partial \mathcal{H}}{\partial q_{i,y}} \rangle + \langle q_{i,z} \frac{\partial \mathcal{H}}{\partial q_{i,z}} \rangle = 3k_B T$

$= \langle \vec{q}_i \cdot \frac{\partial \mathcal{H}}{\partial \vec{q}_i} \rangle = - \langle \vec{q}_i \cdot \vec{F}_i \rangle$

$\Rightarrow \sum_{i=1}^N \langle \vec{q}_i \cdot \vec{F}_i \rangle = -3Nk_B T$ with $\vec{F}_i$ the total force applied to part i

proof of the theorem

$\langle x_i \frac{\partial \mathcal{H}}{\partial x_j} \rangle = \frac{\int d\Gamma x_i \frac{\partial \mathcal{H}}{\partial x_j} e^{-\beta \mathcal{H}}}{\int d\Gamma e^{-\beta \mathcal{H}}} = \frac{1}{\beta \mathcal{Z}} \int d\Gamma x_i \frac{\partial}{\partial x_j} (e^{-\beta \mathcal{H}})$

$= \frac{1}{\beta \mathcal{Z}} \int d\Gamma \left[\frac{\partial}{\partial x_j} (x_i e^{-\beta \mathcal{H}}) - e^{-\beta \mathcal{H}} \frac{\partial x_i}{\partial x_j} \right]$

$= \frac{1}{\beta \mathcal{Z}} \int d\Gamma e^{-\beta \mathcal{H}} \delta_{ij}$

$= k_B T \delta_{ij}$ $\square$

$\frac{\partial}{\partial x_j} (x_i e^{-\beta \mathcal{H}}) = \frac{\partial x_i}{\partial x_j} e^{-\beta \mathcal{H}} + x_i \frac{\partial e^{-\beta \mathcal{H}}}{\partial x_j}$

$\int d\Gamma \frac{\partial}{\partial x_j} (x_i e^{-\beta \mathcal{H}}) = \int d\Gamma \delta_{ij} (x_i e^{-\beta \mathcal{H}}) |_{x_j=-\infty}^{x_j=+\infty} = 0$

$\int dp_{1,x} dp_{1,y} \dots dp_{N,x} dp_{N,y} \dots dp_{N,z} \left[\frac{\partial}{\partial p_{N,y}} (x_i e^{-\beta \mathcal{H}}) \right]$

partial integration $= \int dp_{1,x} dp_{1,y} \dots dp_{N,x} dp_{N,y} \dots dp_{N,z} \left[x_i e^{-\beta \mathcal{H}} \right]_{p_{N,y}=-\infty}^{p_{N,y}=+\infty} = 0$

with $e^{-\beta \mathcal{H}} \sim e^{-p_{N,y}^2/2m}$

$\Rightarrow 0$

for the potential: $\varphi \rightarrow +\infty$ at boundaries since there is a wall confining the parcel within volume V

Diatomic molecules

$\mathcal{H}_{mol} = \frac{\vec{p}_1^2}{2m} + \frac{\vec{p}_2^2}{2m} + \frac{k}{2} |\vec{q}_1 - \vec{q}_2|^2$

$\mathcal{Z}(N, V, T) = \frac{\mathcal{Z}_{mol}(V, T)^N}{N!}$

$\rightarrow \mathcal{Z}_{mol}(V, T) = \frac{1}{h^6} \int d\vec{p}_1 d\vec{p}_2 d\vec{q}_1 d\vec{q}_2 e^{-\beta [\frac{\vec{p}_1^2}{2m} + \frac{\vec{p}_2^2}{2m} + \frac{k}{2} |\vec{q}_1 - \vec{q}_2|^2]}$

$= \frac{1}{\lambda_T^6} \int d\vec{q}_1 d\vec{q}_2 e^{-\beta k/2 |\vec{q}_1 - \vec{q}_2|^2}$

$= \frac{1}{\lambda_T^6} \int d\vec{q}_{cm} d\vec{q} e^{-\beta k/2 \vec{q}^2}$ $\left. \begin{array}{l} \vec{q} = \vec{q}_1 - \vec{q}_2 \\ \vec{q}_{cm} = \frac{\vec{q}_1 + \vec{q}_2}{2} \end{array} \right\}$

$= \frac{V}{\lambda_T^6} \int d\vec{q} e^{-\beta k \vec{q}^2/2} = \frac{V}{\lambda_T^6} \left(\frac{\pi}{\beta k} \right)^{3/2}$

$\rightarrow F = -k_B T \log \mathcal{Z} = -k_B T \log \frac{\mathcal{Z}_{mol}^N}{N!} \approx -Nk_B T \log \mathcal{Z}_{mol} + k_B T (N \log N - N)$
 $\log N! \approx N \log N - N$
 $= -Nk_B T \log \frac{V}{N \lambda_T^3} \left(\frac{2\pi}{\beta k} \right)^{3/2} - Nk_B T$

$\rightarrow C_V = \frac{\partial E}{\partial T} \Big|_{V,N}$, $E = \langle \mathcal{H} \rangle = -\frac{\partial \log \mathcal{Z}}{\partial \beta}$

$\mathcal{Z}_{mol} \sim \frac{1}{\lambda_T^6} \frac{1}{\beta^{3/2}} \sim (\beta^{1/2})^{-6} \frac{1}{\beta^{3/2}} \sim \frac{1}{\beta^3} \frac{1}{\beta^{3/2}} \sim \frac{1}{\beta^{9/2}} \rightarrow \log \mathcal{Z}_{mol} = \dots + \log \beta^{9/2}$

$\rightarrow E_{mol} = \frac{9}{2} k_B T \rightarrow C_{V,mol} = \frac{9}{2} k_B$

$\rightarrow \langle |\vec{q}_1 - \vec{q}_2|^2 \rangle = \langle \vec{q}^2 \rangle$

$\rightarrow \mathcal{H}_{mol} = \dots + \frac{\vec{q}^2}{2} \rightarrow \text{equip } \langle \vec{q} \cdot \frac{\partial \mathcal{H}}{\partial \vec{q}} \rangle = \langle \vec{q}^2 \rangle = 3k_B T \rightarrow \langle \vec{q}^2 \rangle = \frac{3k_B T}{k}$

and $\langle \vec{q}^2 \rangle = -\frac{2}{\beta} \frac{\partial \log \mathcal{Z}}{\partial k} \rightarrow \mathcal{Z}_{mol} \sim k^{-3/2}$

Energy fluctuations of ideal gases

Describing fluctuations: σ_a for quantity a
 $\frac{\sigma_a^2}{N}$: how they change with N ; $\frac{\sigma_a}{\langle a \rangle}$: relative fluct.

$$\sigma_E^2 = \langle E^2 \rangle - \langle E \rangle^2 = \frac{\partial^2 \log Z}{\partial \beta^2}$$

proof $Z = \int d\Gamma e^{-\beta H(\Gamma)} = \sum e^{-\beta E}$

$$\rightarrow -\frac{1}{Z} \frac{\partial Z}{\partial \beta} = \left(-\frac{\partial \log Z}{\partial \beta} \right) = \langle E \rangle$$

$$\frac{1}{Z} \frac{\partial^2 Z}{\partial \beta^2} = \langle E^2 \rangle$$

$$\frac{\partial^2 \log Z}{\partial \beta^2} = \frac{\partial}{\partial \beta} \left[\frac{1}{Z} \frac{\partial Z}{\partial \beta} \right] = \frac{1}{Z} \frac{\partial^2 Z}{\partial \beta^2} - \frac{1}{Z^2} \left(\frac{\partial Z}{\partial \beta} \right)^2$$

$$= \langle E^2 \rangle - \langle E \rangle^2$$

$$\sigma_N^2 = \langle N^2 \rangle - \langle N \rangle^2 = \frac{\partial^2 \log \Xi}{\partial \mu^2}$$

proof $\Xi = \sum_N e^{\beta \mu N} Z(N, V, T)$

$$\rightarrow + \frac{\partial \log \Xi}{\partial \mu} = \frac{1}{\beta \Xi} \frac{\partial \Xi}{\partial \mu} = \langle N \rangle$$

$$\frac{1}{\beta \Xi} \frac{\partial \Xi}{\partial \mu} = \frac{1}{\beta \Xi} \sum_N \beta N e^{\beta \mu N} Z(N, V, T) = \langle N \rangle$$

$$\frac{1}{\beta^2 \Xi} \frac{\partial^2 \Xi}{\partial \mu^2} = \langle N^2 \rangle$$

$$\frac{1}{\beta^2 \Xi} \frac{\partial^2 \Xi}{\partial \mu^2} = \langle N^2 \rangle$$

$$\frac{\partial^2 \log \Xi}{\partial \mu^2} = \frac{1}{\beta} \frac{\partial}{\partial \mu} \left[\frac{1}{\beta \Xi} \frac{\partial \Xi}{\partial \mu} \right] = \frac{1}{\beta^2 \Xi} \frac{\partial^2 \Xi}{\partial \mu^2} - \frac{1}{\beta^2 \Xi^2} \left(\frac{\partial \Xi}{\partial \mu} \right)^2 = \langle N^2 \rangle - \langle N \rangle^2$$

→ for an ideal gas

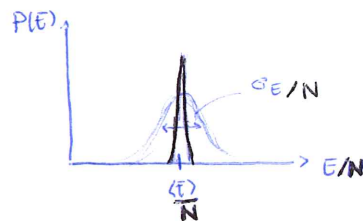
$$Z = \frac{V^N}{N! \lambda^3} \sim \beta^{-3N/2}$$

$$\Rightarrow \log Z = \dots - \frac{3N}{2} \log \beta$$

$$\rightarrow \frac{\partial^2 \log Z}{\partial \beta^2} = \frac{3N}{2} \frac{1}{\beta^2} \Rightarrow \sigma_E^2 = \frac{3N}{2} (k_B T)^2$$

$$\text{and } \langle E \rangle = \frac{3N k_B T}{2}$$

$$\rightarrow \frac{\sigma_E^2}{N} \sim \frac{1}{N}$$



→ N larger $\Rightarrow$ fluctuations smaller and smaller

$$N \rightarrow \infty : \frac{\sigma_E}{\langle E \rangle} \rightarrow 0$$

$$\frac{\sqrt{\frac{3N}{2}} (k_B T)}{\frac{3N}{2} k_B T} = \frac{1}{\sqrt{3N/2}}$$

→ ensembles become equivalent, canonical $\rightarrow$ microcanonical

Harmonic oscillator in 1D

$$\psi(q) = \frac{k}{2} q^2$$

$$\rightarrow Z = \dots \int dq e^{-\beta \frac{k}{2} q^2}$$

$$= \dots \frac{1}{\sqrt{\beta k}} \int = \dots \frac{1}{\beta^{1/2}} \Rightarrow \text{only look at contribution oscillation} \rightarrow E = \frac{k_B T}{2}$$

for 2 oscillations $\frac{k_1}{2} q_1^2 + \frac{k_2}{2} q_2^2$

$$Z = \dots \frac{1}{\sqrt{\beta k_1}} \frac{1}{\sqrt{\beta k_2}} \int \dots \rightarrow \langle E \rangle = \frac{k_B T}{2} + \frac{k_B T}{2}$$

* EQUIPARTITION THEOREM

every quadratic degree of freedom contributes as $\frac{k_B T}{2}$

$$\Rightarrow \text{for } H_{\text{mol}} = \frac{\vec{p}_1^2}{2m} + \frac{\vec{p}_2^2}{2m} + \frac{k}{2} |\vec{q}_1 - \vec{q}_2|^2$$

3 quadr. dof $\rightarrow 3$

$$\Rightarrow \frac{9 k_B T}{2}$$

careful for situations like these:

$$\psi(q) = \frac{k}{2} q^2 + \frac{k'}{4} q^4 \rightarrow \text{can't apply equip.}$$

$$\text{but } \psi(q) = \frac{k'}{4} q^4 \rightarrow \text{equip: } E = \frac{k_B T}{4}$$

Recap

$$p_{mc}(\Gamma) = \dots \int (E - H(\Gamma))$$

E, N, V

E_{fixed}

$$p_c(\Gamma) = \dots e^{-\beta H(\Gamma)}$$

T, N, V

T_{fixed}

- "microcanonical" ensemble in CSFT

$$p_{gc}(\Gamma) = \dots e^{-\beta H(\Gamma)} e^{\beta \mu N}$$

T, μ, V

T_{fixed}

- grand ensemble in CSFT

3 Interacting Systems

so far we have used $Z(N, V, T) = \frac{1}{N!} [Z(1, V, T)]^N$
however this is NOT valid if we have interactions! (interactions $\rightarrow$ potentials)

Pair potentials

$$K(\Gamma) = \sum_{i=1}^N \vec{p}_i^2 / 2m + \Phi(\vec{q}_1, \dots, \vec{q}_N)$$

$\rightarrow$ pairwise interactions: $\Phi = \sum_{i,j} \psi(|\vec{q}_i - \vec{q}_j|)$
thus $Q(N, V, T) = \int d\vec{q}_1 \dots d\vec{q}_N e^{-\beta \sum_{i,j} \psi_{ij}}$
 $= \int d\vec{q}_1 \dots d\vec{q}_N e^{-\beta \psi_{12}} e^{-\beta \psi_{13}} \dots e^{-\beta \psi_{23}} \dots e^{-\beta \psi_{N-1,N}}$

example Lennard-Jones potential



$$\psi_{LJ}(r) = \epsilon \left[\left(\frac{\sigma}{r} \right)^{12} - \left(\frac{\sigma}{r} \right)^6 \right]$$

repulsive at short distances $\leftarrow$ strong repulsive energy against overlapping of s -orbitals
attractive at longer distances $\leftarrow$ van der Waals interactions between neutral atoms

Hard Sphere potential



$$\psi_{HS}(r) = \begin{cases} +\infty & 0 < r < \sigma \\ 0 & r \geq \sigma \end{cases}$$

$\rightarrow$ short distance repulsion (only)

$\rightarrow$ since $Z(N, V, T) = \frac{1}{N! \lambda_1^{3N}} \int d\vec{q}_1 \dots d\vec{q}_N e^{-\beta \sum_{i,j} \psi_{ij}}$ cannot be computed as Φ cannot be factorized
 $\Rightarrow$ approximated techniques useful if SYSTEM IS DILUTED (n low)
should include hard wall pot.

$\rightarrow$ virial expansion: P as an expansion in powers of $n = N/V$

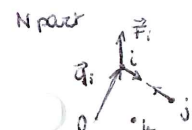
$$P = nk_B T (1 + b_2 n + b_3 n^2 + \dots)$$

ideal gas mutual interactions between particles

Virial theorem

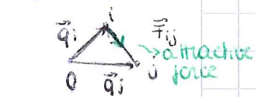
$$\vec{F}_i = \vec{F}_i^{(w)} + \sum_{j \neq i} \vec{F}_{ij}$$

$\rightarrow$ total force = force from collisions with the wall of the container + force from intermolecular interactions
force that j exerts on i



$$\frac{\partial \psi_{ij}}{\partial q_i} = \frac{\partial \psi_{ij}}{\partial q_j} \frac{q_{j,i} - q_{i,j}}{q_{ij}}$$

$$|\vec{q}_i - \vec{q}_j| = q_{ij}$$



$$\rightarrow \text{equip: } \sum_{i=1}^N \langle \vec{q}_i \cdot \vec{F}_i^{(w)} \rangle + \sum_{i,j=1}^N \langle \vec{q}_i \cdot \vec{F}_{ij} \rangle = -3Nk_B T \quad \text{with } \vec{F}_{ii} = 0$$

$$+ \vec{F}^{(w)} = \frac{\Delta \vec{p}}{\Delta t} = -\frac{2\vec{p}_\perp}{\Delta t}$$

force perpendicular to wall, pointing inwards

particles collide elastically with the wall

now $\vec{p} = \vec{p}_\parallel + \vec{p}_\perp$ ($\parallel$ to wall, $\perp$ to wall momentum)

$$\vec{p}' = -\vec{p}_\perp + \vec{p}_\parallel \quad \text{after collision}$$

$$\Delta \vec{p} = \vec{p}' - \vec{p} = \vec{p}_\parallel - \vec{p}_\perp - \vec{p}_\perp - \vec{p}_\parallel = -2\vec{p}_\perp$$

$\vec{p}$ = pressure = average force per unit area exerted on the wall by the particles

with $\hat{n}$ perpendicular in $\vec{x}$, pointing outwards

$$\rightarrow \sum_{i=1}^N \langle \vec{q}_i \cdot \vec{F}_i^{(w)} \rangle = -P \oint \vec{r} \cdot d\vec{S} = -P \int \vec{r} \cdot \vec{e} dV = -3P \int dV = -3PV$$

$$+ \sum_{i,j=1}^N \langle \vec{q}_i \cdot \vec{F}_{ij} \rangle = \frac{1}{2} \sum_{i,j=1}^N \langle \vec{q}_i \cdot \vec{F}_{ij} + \vec{q}_j \cdot \vec{F}_{ji} \rangle = \frac{1}{2} \sum_{i,j=1}^N \langle (\vec{q}_i - \vec{q}_j) \cdot \vec{F}_{ij} \rangle \quad (\vec{F}_{ij} = -\vec{F}_{ji}, N \text{ III})$$

$$= -\frac{1}{2} \sum_{i,j=1}^N \langle q_{ij} \frac{d\psi(q_{ij})}{dq_{ij}} \rangle$$

$$\vec{F}_{ij} = -\frac{d\psi_{ij}}{dq_{ij}} \frac{\vec{q}_i - \vec{q}_j}{q_{ij}}$$

(see back of this page)

$$\sum_{i,j=1}^N \langle \vec{q}_i \cdot \vec{F}_{ij} \rangle = -\frac{1}{2} \sum_{i,j=1}^N \langle q_{ij} \frac{d\psi(q_{ij})}{dq_{ij}} \rangle$$

$$1 = \int d\vec{r} d\vec{r}' \delta(\vec{r} - \vec{q}_i) \delta(\vec{r}' - \vec{q}_j) = \int d\vec{r} d\vec{r}' \delta(\vec{r} - \vec{q}_i) \delta(\vec{r} + \vec{r}' - \vec{q}_j)$$

$$= -\frac{1}{2} \sum_{i,j=1}^N \langle q_{ij} \frac{d\psi(q_{ij})}{dq_{ij}} \int d\vec{r} d\vec{r}' \delta(\vec{r} - \vec{q}_i) \delta(\vec{r} + \vec{r}' - \vec{q}_j) \rangle$$

$$= -\frac{1}{2} \int d\vec{r} d\vec{r}' \rho \frac{d\psi(\rho)}{d\rho} \left(\sum_{i=1}^N \delta(\vec{r} - \vec{q}_i) \sum_{j=1}^N \delta(\vec{r} + \vec{r}' - \vec{q}_j) \right)$$

$$= -\frac{1}{2} \int d\vec{r} d\vec{r}' \rho \frac{d\psi(\rho)}{d\rho} n^{(2)}(\vec{r}, \vec{r} + \vec{r}')$$

with $n^{(2)}(\vec{r}, \vec{r} + \vec{r}')$ the pair correlation function

= prob of finding 2 particles in the volume elements $d\vec{r}$ and $d\vec{r}'$ around $\vec{r}$ and $\vec{r}'$ ($\vec{r}' = \vec{r} + \vec{r}'$)

→ for a homogeneous system, far away from walls, with transl. & rot invariance:

for a system

$$n^{(2)}(\vec{r}, \vec{r} + \vec{r}') = n^2 g(\rho)$$

with $n = N/V$, $g(\rho) =$ radial correlation function

for $\rho \rightarrow \infty$ (large distances) → prob. of finding the 2 part. factorizes into correlation

$$n^{(2)} \rightarrow n \rightarrow g(\rho) \rightarrow 1$$

$$= -\frac{N^2}{2V} \int \rho \frac{d\psi(\rho)}{d\rho} g(\rho) d\vec{r}$$

$$\int d\vec{r} n^{(2)}(\vec{r}, \vec{r}') = V n^2 g(\rho)$$

$$\Rightarrow -3Nk_B T = -3PV - \frac{N^2}{2V} \int \rho \frac{d\psi(\rho)}{d\rho} g(\rho) d\vec{r}$$

$$\Rightarrow P = nk_B T - \frac{n^2}{6} \int \rho \frac{d\psi(\rho)}{d\rho} g(\rho) d\vec{r}$$

→ for non-interacting particles $\psi = 0$
⇒ reduces to ideal gas law

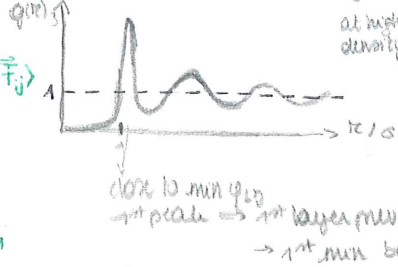
$$n^2 g(\rho)$$

$$= \sum_{i,j} \langle \delta(\vec{r} - \vec{q}_i) \delta(\vec{r} + \vec{r}' - \vec{q}_j) \rangle$$

! EXACT EXPRESSION

given by an integral over the positions of all particles
in general not solvable

example $g(r)$: Lennard-Jones



vanishes at short distances - due to repulsive part ψ_{LD}

at high ρ → damped oscillations

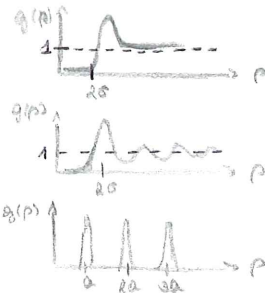
→ prefer particles to be found @ specific distances from the origin

lower ρ → oscillations reduced

1st peak remains due to attraction, but $g(r) \rightarrow 1$ much more rapidly

for ideal gas: $g(\rho) = 1 \rightarrow g(r)$ to this value

other example



for hard spheres



→ crystal structure



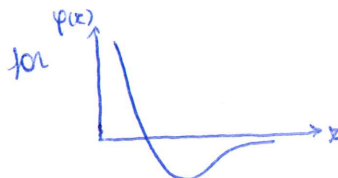
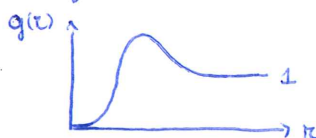
Virial expansion

radial distribution function $g(r)$ describes how the probability of finding another particle varies with r (given a fixed ref. part at $r=0$)

! assume DILUTE system: very low density

→ assume only one particle in its surroundings

$$\Rightarrow g(r) \approx e^{-\beta\psi(r)}$$



assume only 1 pair gets close



→ for $r \rightarrow \infty$: $g(r) \rightarrow 1$ as $\psi(r) \rightarrow 0$

neglect correlations between particles beyond pair interactions

oscillations ψ_{ij} not possible w/o pair correlation since it manifests itself only at high densities

integration by parts
 $\int_a^b u v' dx = [uv]_a^b - \int_a^b u' v dx$

low n

for $g(r) \approx e^{-\beta\phi(r)}$, we get

$$P = nk_B T - \frac{n^2}{6} \int r \frac{d\phi(r)}{dr} g(r) dr = nk_B T - \frac{n^2}{6} \int r \frac{d\phi(r)}{dr} e^{-\beta\phi(r)} dr$$

$$= nk_B T + \frac{n^2}{6} \int r \cdot \frac{d}{dr} (e^{-\beta\phi(r)} - 1) dr$$

$$= nk_B T - \frac{n^2 k_B T}{2} \int (e^{-\beta\phi(r)} - 1) dr$$

$$= nk_B T (1 + b_2 n)$$

$\hookrightarrow$ 1st correction to ideal gas law

$$\text{with } b_2 = \frac{1}{2} \int f(r) dr$$

2nd virial coefficient

$$\text{and } f(r) = e^{-\beta\phi(r)} - 1$$

the Mayer f-function
temp. dependent

@ low T: f has a strong peak
interpart. attraction dominates

$$\Rightarrow b_2 < 0, T_1$$

$$\rightarrow P < P_{\text{ideal gas}}$$

$$@ T = T_{\text{Boyle}} : b_2 = 0$$

$\rightarrow$ behaviour next close to ideal gas - corrections of higher order in n

@ high T: attractive part inter. negligible

$\rightarrow$ dominated by short dist. repulsion

$$\Rightarrow b_2 > 0, T_2$$

$$\rightarrow P > P_{\text{ideal gas}}$$

correction $O(n^2)$

insistence no effect and useful later
will avoid divergences at boundaries

integration by parts with change of coord

$$\begin{aligned} dr &\rightarrow 4\pi r^2 dr \\ &\Rightarrow -\frac{4\pi}{3} \int_0^\infty r^3 \frac{d}{dr} (e^{-\beta\phi(r)} - 1) dr \\ &= -\frac{4\pi}{3} \left[r^3 (e^{-\beta\phi(r)} - 1) \right]_0^\infty + \int_0^\infty 3r^2 (e^{-\beta\phi(r)} - 1) dr \end{aligned}$$

since if $\phi(r) \sim \frac{A}{r^6}$ (LJ)

$$\rightarrow r^3 (e^{-\beta A/r^6} - 1) \sim \frac{A}{r^3} \rightarrow 0$$

virial expansion:

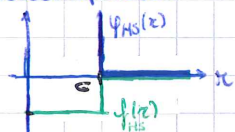
$$P = nk_B T (1 + b_2 n + b_3 n^2 + \dots)$$

with b_k the k^{th} virial coeff

$$\text{and } b_3 = -\frac{1}{3} \int dr_1 dr_2 dr_3 f(r_1) f(r_2) f(r_1 - r_2)$$

examples:

Hard spheres

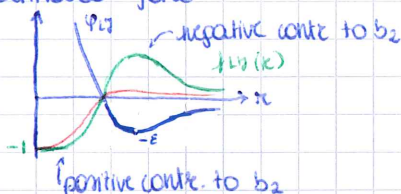


$$\phi_{HS} = \begin{cases} 0 & r \geq \sigma \\ +\infty & r < \sigma \end{cases}$$

$$\rightarrow b_2 = -\frac{1}{2} \int_0^\infty dr 4\pi r^2 (e^{-\beta\phi_{HS}} - 1) = \frac{2\pi\sigma^3}{3} > 0, \text{ indep. on } T$$

$$\rightarrow \text{hard spheres are athermal}, \quad Z(N,V,T) = \frac{Q(N,V,T)}{N! \lambda_T^{3N}}$$

Lennard-Jones



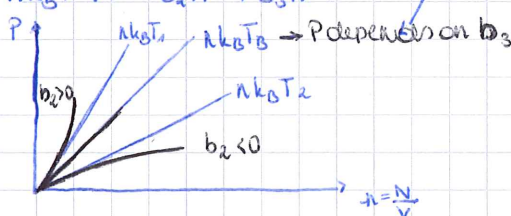
if temp is high: $\beta\phi \ll 1, \beta\epsilon \ll 1$

$\rightarrow b_2$ expected to be positive

finding Boyle temp.

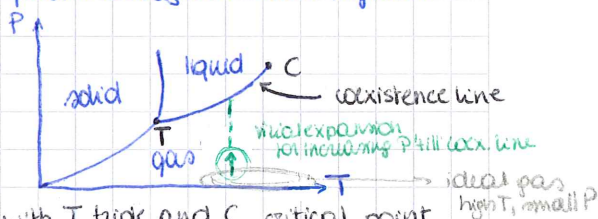


$$P = nk_B T (1 + b_2 n + b_3 n^2 + \dots)$$



$\hookrightarrow$ look @ the virial exponents to see how a gas deviates from the ideal gas

phase diagram real system



with T triple and C critical point

two phase can coexist if

$$P_A = P_B \quad \times \quad T_A = T_B \quad \times \quad \mu_A = \mu_B$$

mechanical equilibrium thermal equilibrium chemical equilibrium

The van der Waals model

$$P = \frac{Nk_B T}{V - Nb} - \frac{aN^2}{V^2} \quad a, b > 0$$

→ low density limit (n small)

$$P = \frac{Nk_B T}{V - Nb} - \frac{aN^2}{V^2} = \frac{Nk_B T}{V} \left(1 + nb + n^2 b^2 + \dots \right) - \frac{aN^2}{V^2}$$

$$= nk_B T \left[1 + n \left(b - \frac{a}{k_B T} \right) + \dots \right]$$

compare with virial expansion and we find

$$b_2 = b - \frac{a}{k_B T}$$

where thus

b should come from

repulsion (excluded volume interaction)

a should come from attraction (intermolecular)



$$k_B T_B = \frac{a}{b}$$

$$[b_2(T_B) = 0]$$

repulsion
→ $b_2 > 0$
attraction
→ $b_2 < 0$

- Bogoliubov inequality

$$\mathcal{H} = \mathcal{H}_0 + \mathcal{H}_1 \quad \text{interaction terms}$$

$$\mathcal{Z} = \sum e^{-\beta \mathcal{H}} = \sum e^{-\beta \mathcal{H}_0} e^{-\beta \mathcal{H}_1} = \mathcal{Z}_0 \frac{\sum e^{-\beta \mathcal{H}_0} e^{-\beta \mathcal{H}_1}}{\sum e^{-\beta \mathcal{H}_0}} = \mathcal{Z}_0 \langle e^{-\beta \mathcal{H}_1} \rangle_0$$

$$\text{now } \langle e^{-\beta \mathcal{H}_1} \rangle_0 \geq e^{-\beta \langle \mathcal{H}_1 \rangle_0}$$

proof: for any $x \in \mathbb{R}$: $\exp(x) \geq 1 + x$

now x is a random variable distributed according to some distr. $p(x) \geq 0$

$$\rightarrow \langle f \rangle = \int dx f(x) p(x)$$

$$\Rightarrow \langle e^{x - \langle x \rangle} \rangle \geq \langle 1 + (x - \langle x \rangle) \rangle = 1$$

$$\Rightarrow \langle e^x \rangle \geq e^{\langle x \rangle}$$

$$\text{thus } \mathcal{Z} \geq \mathcal{Z}_0 e^{-\beta \langle \mathcal{H}_1 \rangle_0}$$

$$\text{and } F = -k_B T \log \mathcal{Z} \quad (\text{Helmholtz free energy})$$

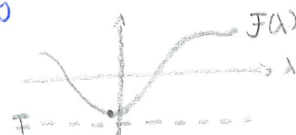
$$\leq -k_B T \log \mathcal{Z}_0 + \langle \mathcal{H}_1 \rangle_0$$

$$\Rightarrow F \leq F_0 + \langle \mathcal{H}_1 \rangle_0 \quad \text{Bogoliubov inequality}$$

$$\text{if } \mathcal{H}(\lambda) = \mathcal{H}_0(\lambda) + \mathcal{H}_1(\lambda)$$

$$\rightarrow F \leq F(\lambda) = F_0(\lambda) + \langle \mathcal{H}_1(\lambda) \rangle_0$$

Bogoliubov: $\hat{F} = \min_{\lambda} F(\lambda)$ is best approx of free energy



- Derivation van der Waals

$$\mathcal{H} = \sum_{i=1}^N \left(\frac{p_i^2}{2m} + \sum_{i < j}^N \varphi(q_{ij}) \right) \quad \text{with } \varphi(r) = \begin{cases} +\infty & r \leq \sigma \\ -\epsilon \left(\frac{\sigma}{r} \right)^6 & r > \sigma \end{cases} \rightarrow \text{hard sphere pot + attractive tail } \sim r^{-6}$$

$$= \left(\sum_{i=1}^N \frac{p_i^2}{2m} + N\lambda \right) + \left(\sum_{i=1}^N \sum_{i < j}^N \varphi(q_{ij}) - N\lambda \right) = \mathcal{H}_0(\lambda) + \mathcal{H}_1(\lambda)$$

$$\Rightarrow \mathcal{Z}_0 = \frac{V^N}{N! \lambda_T^{3N}} e^{-\beta N \lambda}$$

$$\rightarrow F_0(\lambda) = -Nk_B T \log \left(\frac{V}{N \lambda_T^3} \right) - Nk_B T + N\lambda$$

$$\rightarrow \log N! = N \log N - N$$

$$\langle \mathcal{H}_1 \rangle_0 = \frac{\int d\vec{q}_1 \dots d\vec{q}_N \left(\sum_{i=1}^N \sum_{i < j}^N \varphi(q_{ij}) - N\lambda \right) e^{-\beta N \lambda}}{\int d\vec{q}_1 \dots d\vec{q}_N e^{-\beta N \lambda}}$$

$$\sum_{i=1}^N \sum_{i < j}^N \rightarrow \frac{N(N-1)}{2}$$

$$= \frac{1}{V^N} \left\{ \frac{N(N-1)}{2} \int d\vec{q}_1 \dots d\vec{q}_N \varphi(q_{12}) - N\lambda V^N \right\}$$

$$V^{N-2} \int d\vec{q}_1 d\vec{q}_2 \varphi(q_{12})$$

$$= V^{N-1} \int d\vec{q} \varphi(q)$$

$$\vec{q} = \vec{q}_1 - \vec{q}_2$$

$$d\vec{q} = d\vec{q}_1 d\vec{q}_2$$

$$= \frac{N(N-1)}{2V} \int d\vec{q} \varphi(\vec{q}) - N\lambda \approx \frac{N^2}{2V} \int d\vec{q} \varphi(\vec{q}) - N\lambda$$

$$\frac{N}{2} \int d\vec{q} \varphi(\vec{q})$$

$$\rightarrow F(\lambda) = F_0(\lambda) + \langle \mathcal{H}_1(\lambda) \rangle_0 = -Nk_B T \log \left(\frac{V}{N \lambda_T^3} \right) - Nk_B T + \frac{N^2}{2V} \int d\vec{q} \varphi(\vec{q})$$

free energy approx. of interacting system

ideal gas

corrections due to interactions

taken into account with a spatial average

here F independent of λ , no minimum necessary
 $\rightarrow$ pressure $P = -\frac{\partial F}{\partial V} \approx -\frac{\partial F}{\partial V} = \frac{Nk_B T}{V} + \frac{N^2}{2V^2} \int d\vec{q} \varphi(\vec{q})$

note how the integral diverges: $\varphi = \begin{cases} +\infty & r < \sigma \\ -1/r^6 & r > \sigma \end{cases}$ (1)
 (2)



$\Rightarrow$ split interaction pot. into 2 parts

(1) hard sphere repulsion

excluded volume
 $\rightarrow$ effective volume available: $V - Nb$

$\rightarrow$ correction ideal gas law: $V \rightarrow V - Nb$

(2) attractive tail

$$\Rightarrow P \approx \frac{Nk_B T}{V - Nb} + \frac{N^2}{2V^2} \int_{r>\sigma} d\vec{r} \varphi(r)$$

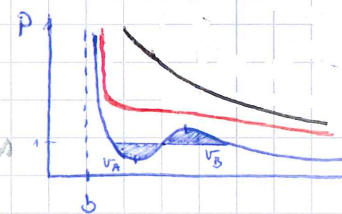
$$= \frac{Nk_B T}{V - Nb} - \frac{aN^2}{V^2}$$

$$a = -\frac{1}{2} \int_{r>\sigma} d\vec{r} \varphi(r)$$

$\rightarrow$ van der Waals!

- Phase separation

below critical point
 $\rightarrow$ separation into 2 different phases



$T > T_c \rightarrow \frac{aN^2}{V^2}$ neglected, pressure decreases monotonically with V

$T = T_c$
 $T < T_c \rightarrow$ for some values of V : $\frac{\partial P}{\partial V} \Big|_{N,T} > 0 \rightarrow$ thermodynamic instability

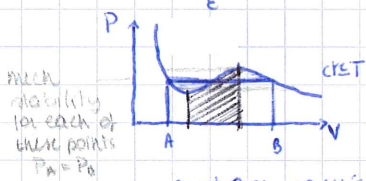
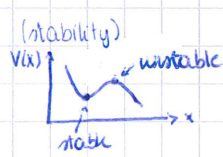
thermodyn. instability $\frac{\partial P}{\partial V} > 0$:



i) $\frac{\partial P}{\partial V} < 0$: P_A decreases, P_B increases

ii) $\frac{\partial P}{\partial V} > 0$: unstable equilibrium
 $\rightarrow$ wall will move further until new stable equil. is reached

$\rightarrow$ phase separation: a dense (liquid) phase & a dilute (gas) phase



2 phase coexistence

OK! $\begin{cases} \mu_A = \mu_B \\ P_A = P_B = P_s \\ T_A = T_B \end{cases}$

$$\mu N = E - TS + PV = F + PV = G$$

$$\text{thus } \mu_A + P_A V_A = \mu_B + P_B V_B$$

$$\Rightarrow \mu_A - \mu_B = P_s (V_B - V_A)$$

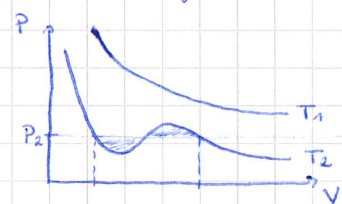
$$= - \int_{V_A}^{V_B} \frac{\partial F}{\partial V} dV$$

$$\text{and since } P(V) = - \frac{\partial F}{\partial V} \Big|_{T,N} = \frac{\partial F}{\partial V} \Big|_{T,N}$$

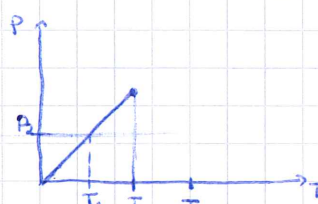
$$\Rightarrow \int_{V_A}^{V_B} P(V) dV = P_s (V_B - V_A)$$

Maxwell construction (equal area law)

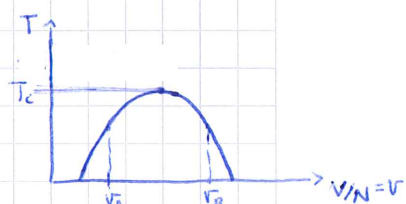
Some diagrams:



$P(N, V, T)$



critical point = endpoint phase coexistence



- Critical point and critical exponents

VDW $P = \frac{Nk_B T}{V - Nb} - \frac{aN^2}{V^2}$ (1)

in vicinity critical point: $\frac{\partial P}{\partial V}|_{N,T} = -\frac{Nk_B T}{(V - Nb)^2} + \frac{2aN^2}{V^3} = 0$ (2)

$\frac{\partial^2 P}{\partial V^2}|_{N,T} = \frac{2Nk_B T}{(V - Nb)^3} - \frac{6aN^2}{V^4} = 0$ (3)

→ from these 3 eq, we find:

$V_c = \frac{V_c}{N} = 3b$, $P_c = \frac{a}{27b^2}$, $k_B T_c = \frac{8a}{27b}$ (*)

→ rescaled thermodynamic variables

$\tilde{P} = P/P_c$, $\tilde{T} = T/T_c$, $\tilde{V} = V/V_c$ $\xrightarrow{\text{VDW}} \tilde{P} \tilde{P}_c = \frac{1}{\tilde{V} \tilde{V}_c - b} - a \frac{1}{\tilde{V}^2 \tilde{V}_c^2}$

$\Rightarrow$ VDW $\tilde{P} = \frac{8\tilde{T}}{3\tilde{V} - 1} - \frac{3}{\tilde{V}^2}$ (mb *)

now in the vicinity of the critical point (@ critical point: $\tilde{P}=1$, $\tilde{T}=1$, $\tilde{V}=1$)

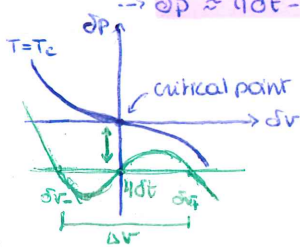
$\tilde{P} = 1 + \delta P$, $\tilde{V} = 1 + \delta V$, $\tilde{T} = 1 + \delta T$ where $\delta P, \delta V, \delta T \ll 1$

→ VDW $(1 + \delta P) = \frac{8(1 + \delta T)}{3(1 + \delta V) - 1} - \frac{3}{(1 + \delta V)^2}$ expand

$\frac{1}{1-x} = (1+x+x^2+\dots)$

$\approx 8(1 + \delta T) \frac{1}{2} (1 - \frac{3}{2}\delta V + \frac{9}{4}\delta V^2 - \frac{27}{8}\delta V^3) - 3(1 - 2\delta V + 3\delta V^2 - 4\delta V^3)$

$= 1 + 4\delta T - 6\delta T\delta V + 9\delta T(\delta V)^2 - \frac{3}{2}(\delta V)^3 + \dots$
 $\Rightarrow \delta P \approx 4\delta T - 6\delta T\delta V + 9\delta T(\delta V)^2 - \frac{3}{2}(\delta V)^3$



1) for $T = T_c \rightarrow \delta T = 0$ along critical isotherm
 $\delta P = -\frac{3}{2}(\delta V)^3$ → power law

2) for $T < T_c \rightarrow \delta T < 0$ (approach T_c from below)

@ $\delta V = 0$: $\delta P = 4\delta T$

$\Rightarrow 4\delta T \approx 4\delta T - 6\delta T\delta V + 9\delta T(\delta V)^2 - \frac{3}{2}(\delta V)^3$

$0 = 6\delta T\delta V - 9\delta T\delta V^2 + 4\delta T = 0$

$\Rightarrow \delta V_{\pm} = \frac{3\delta T \pm \sqrt{9\delta T^2 - 4\delta T}}{3\delta T} \approx \pm \sqrt{-4\delta T} = \pm 2\sqrt{-\delta T}$

maxwell construction $\Delta V = \delta V_+ - \delta V_- = 4\sqrt{-\delta T} = 4(-\delta T)^{1/2}$
 → 2 coexisting phases with $\delta V_+ = \delta V_-$ → power law

2) alternative method:

neglect $9\delta T(\delta V)^2$ since of higher order than $\delta T\delta V$ and $(\delta V)^3$
 $\Rightarrow \delta P = 4\delta T - 6\delta T\delta V - \frac{3}{2}(\delta V)^3$
 now maxwell construction $\delta V_+ = \delta V_-$
 $\Rightarrow \delta P(\delta V_+) = \delta P(\delta V_-)$
 $\Rightarrow 4\delta T - 6\delta T\delta V_+ - \frac{3}{2}(\delta V_+)^3 = 4\delta T - 6\delta T\delta V_- - \frac{3}{2}(\delta V_-)^3$
 $\Rightarrow -12\delta T\delta V_+ = -12\delta T\delta V_- - 3(\delta V_+)^3 + 3(\delta V_-)^3$
 $\Rightarrow \delta V = \delta V_+ - \delta V_- = 2\delta V_+ \sim \sqrt{|\delta T|}$

how the difference ΔV between the volumes of the 2 coex. phases varies as $\delta T \rightarrow 0$

3) isothermal compressibility

$\kappa_T = -\frac{1}{V} \frac{\partial V}{\partial P}|_T$

now $\frac{\partial \tilde{P}}{\partial \tilde{V}} = \frac{-24(1 + \delta T)}{(3\tilde{V} - 1)^2} + \frac{6}{\tilde{V}^3} = -6(1 + \delta T) + 6 = -6\delta T$

$= \frac{\partial P}{\partial V} \frac{V_c}{P_c} = \frac{V_c}{P_c} N \frac{\partial P}{\partial V} = \frac{V_c}{P_c} \frac{\partial P}{\partial V}$

$\Rightarrow \kappa_T \sim \frac{1}{|\delta T|} = (\delta T)^{-1}$ → power

4) specific heat

$E = -\partial_P \log Z = \frac{\partial(\beta F)}{\partial \beta} = \frac{3}{2} Nk_B T - \frac{aN^2}{V}$

$\Rightarrow c_V = \frac{\partial E}{\partial T}|_V = \frac{3Nk_B}{2}$

$\Rightarrow c_V \sim \text{const } (\delta T)^0$ → power

$\Rightarrow \begin{cases} \delta P \sim (\delta V)^3 \\ \delta V \sim (\delta T)^{1/2} \\ \kappa_T \sim |\delta T|^{-1} \\ c_V \sim |\delta T|^0 \end{cases}$

	α	β	γ	δ
VDW	0	1/2	3	1
fluid exp	0.13	0.33	4.8	1.24

→ fluids @ liquid-vapor points (close to critical point) → performed in diff. fluids, all same value: universality

critical exponents
 (universal: do not dep. on a or b)

→ several quantities vanish or diverge as power laws in the vicinity of the critical point

Ising model

VdW fails to quantitatively predict the right exponents due to the used approx in the model - mean-field approx: interaction betw. part. accounted for in averaged way
 ⇒ model beyond mean field approx needed: $\propto \frac{N^2}{2V} \int d\vec{q} \phi(\vec{q})$

Ising model

lattice model with a spin at each lattice point
 $s_i = \pm 1$ → for N spins: 2^N possible configurations

configuration:

$$\mathcal{H} = -J \sum_{\langle i,j \rangle} s_i s_j - H \sum_i s_i$$

with $J > 0$ (ferromagn. case) the coupling strength between spins (interaction)
 $H > 0$ the magnetic field ($H=0$: minimal energy config, all spins "aligned")

↳ pos spin config have lower energies if $H > 0$; neg spins if $H < 0$

$$Z = \sum_{\{s_i\}} e^{-\beta \mathcal{H}(\{s_i\})}$$

→ sum over 2^N states

$$\rightarrow \text{average value of a spin: } \langle s_k \rangle = \frac{1}{Z} \sum_{\{s_i\}} e^{-\beta \mathcal{H}(\{s_i\})} s_k$$

$$\text{magnetization: } M = \sum_k \langle s_k \rangle$$

↳ if $H \neq 0$ → spins predominantly aligned along direction H

$$\rightarrow M \neq 0$$

spontaneous magnetization: magn $M \neq 0$ in absence of magn field ($H=0$)

$$= \lim_{H \rightarrow 0^+} m(T, H) = \pm m_0(T) \neq 0$$

with $m = M/N$

⇒ phase transitions occur

@ coex. line: 2 diff. phases with opposite spontaneous magn

- Absence of phase transition in one dimension

1D:

$$\mathcal{H} = \sum_{i=1}^N \sum_{s_i = \pm 1} \sum_{s_{i+1} = \pm 1} e^{\beta (J \sum_{i=1}^N s_i s_{i+1} + H \sum_{i=1}^N s_i)}$$

$$= T_N (T^N)$$

$$= \lambda_+^N + \lambda_-^N \approx \lambda_+^N$$

T is the transfer matrix: $T = \begin{pmatrix} e^{\beta J - \beta H} & e^{-\beta J - \beta H} \\ e^{-\beta J + \beta H} & e^{\beta J + \beta H} \end{pmatrix}$

$$\rightarrow \lambda_{\pm} = e^{\beta J} \frac{\cosh(\beta H) \pm \sqrt{\cosh^2(\beta H) - \sinh^2(\beta J)}}{\sinh(\beta J)}$$

$$\rightarrow F = -N k_B T \log \lambda_+$$

in limit $N \rightarrow \infty$, largest λ dominates

however no spontaneous magnetization at any finite temp

⇒ one-dim systems with short range interactions have no phase transitions at any finite temperature

however, at zero temp

$$\lim_{H \rightarrow 0} \lim_{T \rightarrow 0^+} m(T, H) = \pm 1$$

→ spont. magn.

$$\Rightarrow T_c = 0 \text{ for 1D Ising}$$

- Two dimensional Ising model: exact solution

spontaneous magn for $T < T_c$:

$$m_0(T) = \left[1 - \sinh^{-4} \left(\frac{2J}{k_B T} \right) \right]^{1/8}$$

with T_c derivable from

$$\sinh \left(\frac{2J}{k_B T_c} \right) = 1$$

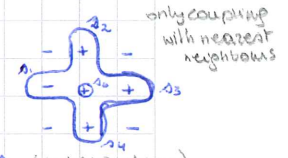
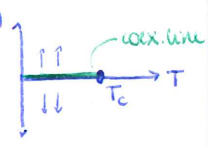
in the vicinity of the critical point: $m_0(T) \sim (T_c - T)^{1/8}$

$$\Rightarrow \beta = 1/8 \text{ since } m_0(T) \sim \delta v \text{ VdW}$$

$$\begin{aligned} M &= \sum_k \langle s_k \rangle \\ &= \sum_k \sum_{\{s_i\}} e^{-\beta \mathcal{H}(\{s_i\})} s_k \\ &= \frac{\partial \log Z}{\partial \beta H} \\ &= - \frac{\partial (-N k_B T \log Z)}{\partial H} \\ &= - \frac{\partial F}{\partial H} \end{aligned}$$

↳ spont. magn: $H=0$

("magnetic model")



$$\lambda_{\pm} = e^{\beta J} \frac{\cosh(\beta H) \pm \sqrt{\cosh^2(\beta H) - \sinh^2(\beta J)}}{\sinh(\beta J)}$$

TODO

2D and 3D do have spont. magn. there is ferromagn.

- Mean-field approximation

approximate solution Ising model using Bogoliubov

$$\rightarrow \mathcal{H} = \mathcal{H}_0(\lambda) + \mathcal{H}_1(\lambda)$$

$$= \underbrace{-\lambda \sum_i s_i}_{\text{non-int}} - \underbrace{J \sum_{\langle ij \rangle} s_i s_j}_{\text{interacting}} - (H-\lambda) \sum_i s_i$$

$$\mathcal{Z}_0 = \sum_{s_i = \pm 1} e^{\beta \lambda \sum_i s_i} = \sum_{s_1 = \pm 1} \sum_{s_2 = \pm 1} \dots \sum_{s_N = \pm 1} e^{\beta \lambda s_1} e^{\beta \lambda s_2} \dots e^{\beta \lambda s_N}$$

$$= \left(\sum_{s = \pm 1} e^{\beta \lambda s} \right)^N = (e^{\beta \lambda} + e^{-\beta \lambda})^N = (2 \cosh(\beta \lambda))^N$$

$$\rightarrow F_0 = -k_B T \log \mathcal{Z}_0 = -N k_B T \log (2 \cosh(\beta \lambda))$$

$$\langle \mathcal{H}_1(\lambda) \rangle_0 = \frac{1}{\mathcal{Z}_0} \sum_{s_i = \pm 1} \left\{ -J \sum_{\langle ij \rangle} s_i s_j - (H-\lambda) \sum_i s_i \right\} e^{-\beta \mathcal{H}_0}$$

they do not interact in $\mathcal{H}_0$

$$= -J \sum_{\langle ij \rangle} \langle s_i s_j \rangle_0 - (H-\lambda) \sum_i \langle s_i \rangle_0 = -\frac{J N z}{2} \langle s \rangle_0^2 - (H-\lambda) N \langle s \rangle_0$$

$$\rightarrow F(\lambda) = F_0(\lambda) + \langle \mathcal{H}_1(\lambda) \rangle_0$$

$$= -N k_B T \log (2 \cosh(\beta \lambda)) - \frac{J N z}{2} \langle s \rangle_0^2 - (H-\lambda) N \langle s \rangle_0$$

$$= -N k_B T \log (2 \cosh(\beta \lambda)) - \frac{J N z}{2} \tanh^2(\beta \lambda) - (H-\lambda) N \tanh(\beta \lambda)$$

$$\rightarrow \hat{F} = \lim_{\lambda} F(\lambda)$$

$$\frac{\partial F}{\partial \lambda} \Big|_{\lambda = \lambda_*} = 0 \Rightarrow -N \tanh(\beta \lambda_*) + N \langle s \rangle - N \frac{\partial \langle s \rangle}{\partial \lambda} (J z \langle s \rangle_0 + H - \lambda) \Big|_{\lambda_*} = 0$$

$$\Rightarrow N \frac{\partial \langle s \rangle}{\partial \lambda} (J z \langle s \rangle_0 + H - \lambda_*) = 0 \Rightarrow J z \langle s \rangle_0 + H - \lambda_* = 0$$

$$\Rightarrow \lambda_* = H + J z \tanh(\beta \lambda_*)$$

$$\hat{F} = F(\lambda_*)$$

$$= -N k_B T \log (2 \cosh(\beta \lambda_*)) - \frac{N J z}{2} \tanh^2(\beta \lambda_*) - N (H - \lambda_*) \tanh(\beta \lambda_*)$$

$$m = - \frac{\partial \hat{F}}{N \partial H} = \frac{-1}{N} \left[\frac{\partial \hat{F}}{\partial \lambda_*} \frac{\partial \lambda_*}{\partial H} + \frac{\partial \hat{F}}{\partial H} \Big|_{\lambda_*} \right]_T$$

0 since λ_* minimizes $\hat{F}$

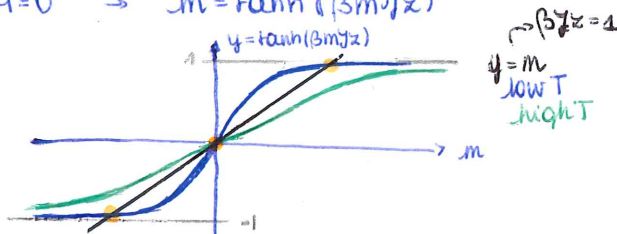
$$= \tanh(\beta \lambda_*) = \tanh[\beta (m J z + H)]$$

$$m = \tanh(\beta \lambda_*)$$

$$\Rightarrow \tanh[\beta (H + J z \tanh(\beta \lambda_*))] = \tanh(\beta (H + J z m))$$

now spontaneous magnetization?

$$H=0 \rightarrow m = \tanh(\beta m J z)$$



vertical point

$$\frac{d \tanh(\beta m J z)}{dm} \Big|_{m=0} = 1$$

$$\tanh(\beta m J z) \approx \beta m J z - \frac{1}{3} (\beta m J z)^3$$

at "high" T, low β ($\beta J z < 1$)

only 1 solution: $m=0$

at sufficiently low T, high β ($\beta J z > 1$)

spont. magn is possible:

$m = \tanh(\beta m J z)$ also has

2 additional solutions: $m = \pm m_0 \neq 0$

$$\beta J z - (\beta m J z)^2 \Big|_{m=0} = \beta_c J z = 1$$

$$\Rightarrow k_B T_c = J z$$

approximation becomes better at higher dimensions

we've approx. the effect of spin-spin interactions with an average field
a given spin s_i appears in the Hamiltonian with a term of type

$$(\sum_j' s_j + H) s_i$$

where the sum is over all its neighbours

if each spin has many neighbours, the approx $\sum_j' s_j \approx \langle s \rangle$ becomes more and more accurate

Critical exponents

several quantities vanish with power-laws in the vicinity of the critical point
 → critical exponents

spontaneous magnetization

$$m_0(T) \sim (T_c - T)^{\beta} \quad \text{along coex line } H=0$$

$$C \sim |T - T_c|^{-\alpha}$$

$$\chi = \frac{\partial m}{\partial H} \sim |T - T_c|^{-\gamma}$$

→ magn. susceptibility

$$H \sim |M|^{\delta} \quad @ T=T_c$$

	α	β	γ	δ	
Ising					
2d	0	1/8	7/4	15	
3d	0.13	0.33	1.24	4.8	→ identical to expm. values for fluids
Bardeen-Gruber mean field	0	1/2	1	3	
→ $m = \tanh[\dots]$					
→ name as mean field VDW					
→ BMF exponents do not dep. on dim.					

correspondence VDW ↔ Ising model:

$$\begin{array}{l} P \leftrightarrow H \\ V \leftrightarrow M \end{array} \quad \text{dim} \sim z$$

* $H \sim m^3 \quad @ T=T_c$

$$h_B T_c = y z \rightarrow m = \tanh(m + \beta_c H)$$

$$\Rightarrow m \approx m + \beta_c H - \frac{1}{3} (m + \beta_c H)^3$$

$$\Rightarrow m^3 \sim H$$

* correlation length ξ = char. distance at which two spins are correlated
 = measure of the char. length scale of fluctuations of the syst.

$$@ T=T_c : \xi \sim |T_c - T|^{-\nu}$$

→ no longer a char. length scale,

system is scale invariant

fluctuations occur at all length scales

the correlation function

$$G^{(2)} \sim \frac{1}{r^{d-2+\eta}}$$

correlation function $G^{(2)}$

for $T > T_c$, $H=0$ (absence spont. magn):

$$G^{(2)}(\vec{r}, \vec{r}') = \langle s_{\vec{r}} s_{\vec{r}'} \rangle$$

spins close by → aligned (more likely)

far away spins → weakly interacting → more likely uncorrelated

for $T < T_c$: $m_0 \neq 0$, $\langle s \rangle \neq 0$

$$G^{(2)}(\vec{r}, \vec{r}') = \langle (s_{\vec{r}} - \langle s \rangle)(s_{\vec{r}'} - \langle s \rangle) \rangle \quad (\text{subtr. spin average})$$

for large distances both decay exponentially

$$G^{(2)}(\vec{r}, \vec{r}') \sim e^{-\frac{|\vec{r} - \vec{r}'|}{\xi}}$$

critical opalescent in fluids:

in vicinity of critical points fluctuations become of a size comparable to the wavelength of light → light is scattered

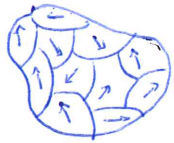
→ normally transparent liquid appears cloudy

→ correlation length ξ diverges at critical point

→ critical behavior indep. on local properties of the Hamiltonian only determined by global properties (dim & symm)

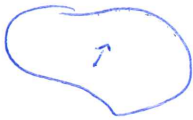
→ many different systems share the same critical behaviour

Magnetism



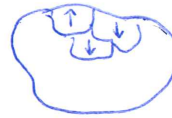
paramagnet

all points magnetized



ferromagnet

all points same magnetization



uniaxial ferromagnet

with thermodynamical variables

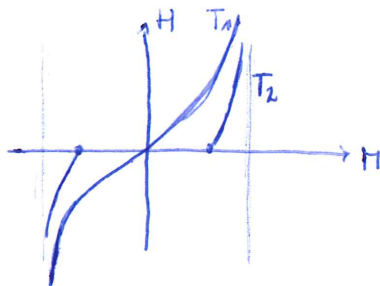
$$M = \sum_i \mu_i$$

magn. moments

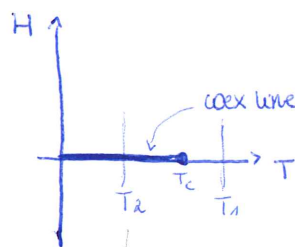
magnetization

H magnetic field (external)

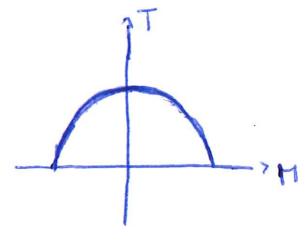
T temperature



at very large H
magnetization $\rightarrow$ asymptotes



coexistence
between 2 phases



4 Quantum Statistical Mechanics

Below: $p(\Gamma) = p(P, q)$ the prob. distribution for 6N-dim config.

Now: + Heisenberg's uncertainty principle

cannot specify state with Γ

$$\Delta p_x \Delta x \geq \hbar/2$$

(no simultaneous info about mom & pos)

$\rightarrow p(P, q)$ (prob. dist.) can't be

All state given by $|\Psi_\alpha\rangle$; 1 part w max prob $\sim \Psi(\vec{r}, t)$

$$\rightarrow \text{SE: } i\hbar \partial_t \Psi(\vec{r}, t) = \hat{H} \Psi(\vec{r}, t)$$

$$\text{with } \hat{H} = -\hbar^2/2m \nabla^2 + V(\vec{r}, t)$$

$\rightarrow |\Psi_\alpha\rangle$ eigenstates $\hat{H}$ with energy E_α

$\rightarrow |\Psi(\vec{q}, T)|^2$: prob. to find particle at time T, position $\vec{q}$

$\Rightarrow$ All states & probabilities

for system in state $|\Psi_\alpha\rangle$

+ All probability

for operator $\hat{A}$ (physical observable):

$$\langle \hat{A} \rangle_{\text{pure}} = \int \Psi_\alpha^*(q) \hat{A} \Psi_\alpha(q) d^{3N}q = \langle \Psi_\alpha | \hat{A} | \Psi_\alpha \rangle$$

+ Quantum statist. mech. prob

$$\langle \hat{A} \rangle = \sum_\alpha c_\alpha \langle \Psi_\alpha | \hat{A} | \Psi_\alpha \rangle_{\text{pure}}$$

with c_α the prob to find the syst. in state $|\Psi_\alpha\rangle$

$$\text{norm.: } \sum_\alpha c_\alpha = 1$$

Using $\sum_n |n\rangle \langle n| = \mathbb{1}$

[complete set of orthonorm. states]

$$\Rightarrow \langle \hat{A} \rangle = \text{Tr}(\hat{A} \hat{\rho})$$

$$\text{where } \hat{\rho} \equiv \sum_\alpha c_\alpha |\Psi_\alpha\rangle \langle \Psi_\alpha|$$

$$= \sum_n \langle n | \hat{A} \hat{\rho} | n \rangle$$

the density matrix

special case: pure state: $\rho = |\Phi\rangle \langle \Phi|$

(no summ)

$\rightarrow$ Quantum canonical ensemble

$$\text{since } \hat{H} |\Psi_\alpha\rangle = E_\alpha |\Psi_\alpha\rangle$$

$$\text{and } c_\alpha = \frac{1}{Z} e^{-\beta E_\alpha}$$

specifies prob distr of states

$$\text{with } \beta = 1/k_B T$$

$$Z = \sum_\alpha e^{-\beta E_\alpha}$$

[norm factor] = PARTITION FUNCTION

$$\rightarrow \hat{\rho}_{\text{can}} = \exp(-\beta \hat{H}) / \text{Tr}(\exp(-\beta \hat{H}))$$

$$\Rightarrow \hat{\rho}(\Gamma) = \begin{pmatrix} \rho(\Gamma_{11}) & \rho(\Gamma_{12}) & \dots & \rho(\Gamma_{1n}) \\ \vdots & \vdots & \ddots & \vdots \\ \rho(\Gamma_{m1}) & \rho(\Gamma_{m2}) & \dots & \rho(\Gamma_{mn}) \end{pmatrix} \text{ for } \Gamma \text{ diag.}$$

$$\Rightarrow e^{\hat{A}} = \mathbb{1} + \hat{A} + \frac{1}{2} \hat{A} \hat{A} + \frac{1}{3!} \hat{A} \hat{A} \hat{A} + \dots$$

example: Quantum Harm. Osc

$$\hat{H} = -\hbar^2/2m \partial_x^2 + \frac{1}{2} m \omega^2 x^2$$

with Hermite polynomials as Ψ_α , and ω the frequency

$$\Rightarrow E_n = \hbar \omega (n + 1/2)$$

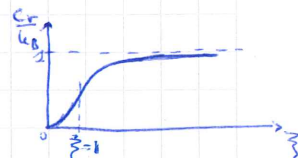
$$n \in \mathbb{N} \quad (0, 1, 2, \dots)$$

(non-degenerate energy spectr.)

$$\rightarrow Z = \sum_{n=0}^{\infty} e^{-\beta E_n} = \frac{e^{-\beta \hbar \omega / 2}}{1 - e^{-\beta \hbar \omega}} = \frac{1}{2 \sinh(\beta \hbar \omega / 2)}$$

$$\rightarrow \langle E \rangle = -\partial_\beta \log Z = \partial_\beta \left[\frac{\beta \hbar \omega}{2} + \log(1 - e^{-\beta \hbar \omega}) \right]$$

$$= \hbar \omega \left[\frac{1}{2} + \frac{1}{e^{\beta \hbar \omega} - 1} \right]$$



$$\text{dimensionful } C_V = \frac{\partial \langle E \rangle}{\partial T} = \frac{k_B (\hbar \omega / k_B T)^2 e^{-\beta \hbar \omega}}{(1 - e^{-\beta \hbar \omega})^2}$$

$$\hat{\xi} = k_B T / \hbar \omega$$

$$\frac{C_V}{k_B} = \frac{1}{\hat{\xi}^2} \cdot \frac{e^{-1/\hat{\xi}}}{(1 - e^{-1/\hat{\xi}})^2}$$

$$\xrightarrow{\text{approx.}} \frac{1 - 1/12 \hat{\xi}^2}{\hat{\xi}^2}$$

$$\hat{\xi} \gg 1$$

high T

purely quantum effect which peaks at some point

$$\hat{\xi} \ll 1$$

low T

$$\Rightarrow \hat{\xi} \gg 1 \rightarrow T \gg \hbar \omega / k_B \text{ around room } \Delta T \rightarrow C_V$$

$$T \gg T_\# \quad C_V$$

$$T \ll T_\# \quad \Delta T$$

$\rightarrow$ charact temp Leading partner in

Test & Mechatronic Simulation

Fermions & Bosons

symm under the exchange of coord of any pair of part.

Q1: particles indistinguishable $\Rightarrow |\Psi(\vec{q}_1, \dots, \vec{q}_i, \dots, \vec{q}_j, \dots, \vec{q}_N)|^2 = |\Psi(\vec{q}_1, \dots, \vec{q}_j, \dots, \vec{q}_i, \dots, \vec{q}_N)|^2$ for $i \neq j$

$\rightarrow$ In $D \neq 2+1$ (2 pos + 1 time) 2 options

$D=2+1 \rightarrow$ anyons
 1) Ψ is symmetric $i \leftrightarrow j \rightarrow$ bosons: $s=0, 1, 2, \dots$ (integer spin) (phase 0)
 2) Ψ is antisymmetric $i \leftrightarrow j \rightarrow$ fermions: $s=1/2, 3/2, \dots$ (half-integer spin) (phase $\pi/2$)

example 2 non-interacting particles

$$H = H_1(\vec{p}_1, \vec{q}_1) + H_2(\vec{p}_2, \vec{q}_2)$$

$$\text{with } H_1 = -\hbar^2/2m \nabla_1^2 + V(\vec{q}_1) \quad \text{and} \quad \hat{H}_1 |\phi_k\rangle = \epsilon_k |\phi_k\rangle$$

k is a quantum number

* quantum state:

$$\Psi(\vec{q}_1, \vec{q}_2) = \phi_k(\vec{q}_1) \phi_l(\vec{q}_2) \quad \text{with } k \neq l$$

$$\begin{aligned} \rightarrow H\Psi &= [H_1(\vec{p}_1, \vec{q}_1) + H_2(\vec{p}_2, \vec{q}_2)] \phi_k(\vec{q}_1) \phi_l(\vec{q}_2) = (H_1(\vec{p}_1, \vec{q}_1) \phi_k(\vec{q}_1)) \phi_l(\vec{q}_2) + \phi_k(\vec{q}_1) (H_2(\vec{p}_2, \vec{q}_2) \phi_l(\vec{q}_2)) \\ &= (\epsilon_k + \epsilon_l) \Psi(\vec{q}_1, \vec{q}_2) \end{aligned}$$

However: no symm: $|\Psi(\vec{q}_1, \vec{q}_2)|^2 \neq |\Psi(\vec{q}_2, \vec{q}_1)|^2$ $k \neq l$

$\Rightarrow$ (anti)symmetrization

$$\Psi_{\pm}(\vec{q}_1, \vec{q}_2) = \frac{1}{\sqrt{2}} [\phi_k(\vec{q}_1) \phi_l(\vec{q}_2) \pm \phi_k(\vec{q}_2) \phi_l(\vec{q}_1)]$$

$$\begin{aligned} \Psi(\vec{q}_2, \vec{q}_1) &= \phi_k(\vec{q}_2) \phi_l(\vec{q}_1) \\ &+ \phi_k(\vec{q}_1) \phi_l(\vec{q}_2) \end{aligned}$$

$\rightarrow \Psi_+$: wavefunction for 2 ident. BOSONS (photons, phonons)

Ψ_- : " " 2 ident. FERMIONS (electrons, protons)

NOW: 2 particles in same quantum states

$$\Psi(\vec{q}_1, \vec{q}_2) = \phi_k(\vec{q}_1) \phi_k(\vec{q}_2)$$

$$\rightarrow \Psi_+ = \sqrt{2} [\phi_k(\vec{q}_1) \phi_k(\vec{q}_2)] \rightarrow \text{Bosons OK}$$

$\Psi_- = 0 \rightarrow$ FERMIONS: PAULI EXCLUSION PRINCIPLE "2 fermions cannot occupy the same quantum state"

N identical particles

n_Y = occupation # (# of particles) in state with ϵ_Y

$\rightarrow$ for example above ($N=2$): $n_k = n_l = 1$, $n_j = 0$ for $j \neq k, l$

general: * bosons: $n_Y \geq 0$

* fermions: $n_Y = 0, 1$ (Pauli)

$$\rightarrow N = \sum_Y n_Y \quad , \quad E = \sum_Y n_Y \epsilon_Y$$

$\rightarrow$ canonical ensemble: partition function

$$\mathcal{Z}(N, V, T) = \sum_{n_1} \sum_{n_2} \dots \sum_{n_Y} e^{-\beta \sum_Y n_Y \epsilon_Y} \underbrace{\delta_{N, \sum_Y n_Y}}_{\substack{\text{constraint: } \sum_Y n_Y = N \\ \text{makes expr. complicated}}}$$

$\rightarrow$ grand canonical ensemble: part. function

$$\begin{aligned} \Xi(\mu, V, T) &= \sum_{N=0}^{\infty} \mathcal{Z}(N, V, T) e^{\beta \mu N} = \sum_{\{n_Y\}} e^{-\beta \sum_Y n_Y \epsilon_Y} \sum_{N=0}^{\infty} e^{\beta \mu N} \delta_{N, \sum_Y n_Y} \\ &= \sum_{\{n_Y\}} e^{-\beta \sum_Y n_Y \epsilon_Y} e^{\beta \mu \sum_Y n_Y} \\ &= \prod_Y \sum_{n_Y} e^{-\beta n_Y (\epsilon_Y - \mu)} \end{aligned}$$

$$e^{a+b} = e^a e^b$$

$$\sum_{n_Y} e^{-\beta n_Y (\epsilon_Y - \mu)} = \sum_{n_Y} e^{-\beta n_Y \epsilon_Y} e^{\beta \mu n_Y}$$

$$\Xi(\mu, V, T) = \prod_r \sum_{n_r} e^{-\beta n_r (\epsilon_r - \mu)}$$

→ Bosons

! convergent only for $e^{\beta(\mu - \epsilon_r)} < 1 \rightarrow \mu < \min(\epsilon_r)$ upper bound μ
 + w/o geom. series ($\sum_{n=0}^{\infty} a r^n = \frac{a}{1-r}$)
 $\Rightarrow \Xi_{BE} = \prod_r \frac{1}{1 - e^{\beta(\mu - \epsilon_r)}}$

→ Fermions

$$\Xi_{FD} = \prod_r [1 + e^{\beta(\mu - \epsilon_r)}] \quad (n_r = 0, 1) \quad \text{no cond. on } \mu$$

$$\Rightarrow \log \Xi = \mp \sum_r \log(1 \mp e^{\beta(\mu - \epsilon_r)})$$

upper sign = bosons
lower sign = fermions

$$\rightarrow \langle n_r \rangle = \frac{-\partial \log \Xi}{\partial \epsilon_r} = \frac{1}{(e^{\beta(\epsilon_r - \mu)} \mp 1)}$$

$\hookrightarrow \exp(\beta(\epsilon_r - \mu)) \gg 1$ bosons $\rightarrow \langle n_r \rangle \gg 0$
 > 0 fermions $\rightarrow \langle n_r \rangle \ll 1$

$$E = \frac{-\partial \log \Xi}{\partial \beta} = \sum_r \frac{\epsilon_r}{e^{\beta(\epsilon_r - \mu)} \mp 1} = \sum_r \langle n_r \rangle \epsilon_r$$

↳ Comments

* FD: $\langle n_r \rangle \neq 0 = \frac{1}{\exp(\beta \mu (\xi - 1)) + 1}$

$$\xi \equiv \epsilon_r / \mu \quad (\text{dimless})$$

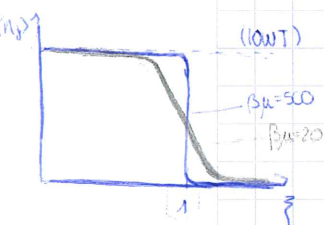
low T ($\beta \rightarrow \infty$): $\langle n_r \rangle = 1$ for $\epsilon_r < \mu$ ($\xi < 1$ → occupied)
 $\sim$ step function $\langle n_r \rangle = 0$ for $\epsilon_r > \mu$ ($\xi > 1$ → free)

→ two temperature limit: $\lim_{T \rightarrow 0} \mu(T) = \epsilon_F$ Fermi-energy

* BEC: $\langle n_r \rangle_{BE} \gg 1$ ($\rightarrow \infty$)

$$\Rightarrow \mu \rightarrow \epsilon_r$$

* Bosons with $\mu=0$: $\langle n_r \rangle = \frac{1}{e^{\beta \epsilon_r} - 1}$ - same as a H₂O



• bosons good approx
 • photons & phonons
 • fermions good for
 electrons in metal

Quantum connections to ideal gas law

(particles in a box)

system of non-interacting, quantum particles in a cubic box of size L
 with single particle energy levels

$$\epsilon_p = \vec{p}^2 / 2m$$

$$\text{SE} \rightarrow -\hbar^2 / 2m \nabla^2 \phi_p = \epsilon_p \phi_p \Rightarrow \phi_{\vec{u}} = e^{i\vec{u} \cdot \vec{x}} \quad (\text{plane wave}) \quad \hbar \vec{u} = \vec{p} = \hbar \vec{k}$$

impose periodic boundary cond $\rightarrow$ quantization $\vec{p} = \hbar \vec{k}$
 $e^{i\vec{k}_x(x+L)} = e^{i\vec{k}_x x} \Rightarrow k_x = \frac{2\pi}{L} n_x \Rightarrow \vec{p} = \frac{\hbar}{L} (n_x, n_y, n_z) \quad n_i \in \mathbb{Z}$

$$\text{now: } \log \Xi = \mp \sum_{\vec{p}} \log(1 \mp e^{\beta(\mu - \vec{p}^2 / 2m)})$$

$$= \mp \frac{V}{h^3} \int d\vec{p} \log(1 \mp \exp(\beta(\mu - \vec{p}^2 / 2m)))$$

$$z = e^{\beta \mu}$$

$$= \mp \frac{V}{h^3} \int d\vec{p} \left[\sum_{l=1}^{\infty} (\pm 1)^l \frac{z^l}{l} e^{-\beta \vec{p}^2 l / 2m} \right]$$

$$= \pm \frac{V}{h^3} \pi^{3/2} (2m k_B T)^{3/2} \sum_{l=1}^{\infty} (\pm 1)^l \frac{z^l}{l^{3/2}}$$

$$= \pm \frac{V}{\lambda_T^3} \sum_{l=1}^{\infty} (\pm 1)^l \frac{z^l}{l^{3/2}}$$

$$\int_{x_1}^{x_2} f(x) dx = \sum_i f(x_i) \Delta x_i$$

$\Delta x = h/L \quad \Delta \vec{p} = \hbar/L$

$$\log(1 \mp z) = - \sum_{l=1}^{\infty} (\pm 1)^l \frac{z^l}{l}$$

$$\int d\vec{p} p^2 e^{-\beta \vec{p}^2 l / 2m} = 4\pi \int_0^{\infty} dp p^2 e^{-\beta p^2 l / 2m} = 4\pi \left(\frac{2m}{\beta l} \right)^{3/2} \sqrt{\pi/4}$$

$$\lambda_T = \frac{h}{\sqrt{2\pi m k_B T}}$$

known:
 $\mu < \epsilon_0 = 0$
 $\beta > 0$
 $\rightarrow z < 1$

$$\log \Xi = \pm \frac{V}{\lambda_T^3} \sum_{l=1}^{\infty} (\pm 1)^l \frac{z^l}{l^{5/2}}$$

$$\rightarrow N = \frac{1}{\beta} \partial_{\mu} \log \Xi = \pm \frac{V}{\lambda_T^3} \sum_{l=1}^{\infty} (\pm 1)^l \frac{z^l}{l^{3/2}}$$

$$\Rightarrow \begin{cases} \frac{PV}{k_B T} = \pm \frac{V}{\lambda_T^3} \sum_{l=1}^{\infty} (\pm 1)^l \frac{z^l}{l^{5/2}} \\ n \lambda_T^3 = \pm \sum_{l=1}^{\infty} (\pm 1)^l \frac{z^l}{l^{3/2}} \end{cases}$$

$$z = e^{\beta \mu} \quad \Rightarrow \quad n = \frac{N}{V} \quad \log \Xi = \frac{PV}{k_B T}$$

~~lowest order~~ expand in z (small)

$$\text{1st order} \quad \begin{cases} \frac{PV}{k_B T} = \frac{V}{\lambda_T^3} z \\ n \lambda_T^3 = z \end{cases} \Rightarrow p = n k_B T \quad \text{classical ideal gas}$$

$$\text{2nd order} \quad \begin{cases} \frac{PV}{k_B T} \approx \frac{V}{\lambda_T^3} \left(z \pm \frac{z^2}{4N^{1/2}} \right) \\ n \lambda_T^3 \approx z \pm \frac{z^2}{2N^{1/2}} \end{cases} \Rightarrow p \approx n k_B T \left(1 \mp \frac{n \lambda_T^3}{4N^{1/2}} \right)$$

$$\text{3rd order} \quad \begin{cases} \frac{PV}{k_B T} = \frac{V}{\lambda_T^3} \left(z \pm \frac{z^2}{4N^{1/2}} + \frac{z^3}{9N^{3/2}} \right) \\ n \lambda_T^3 = z \pm \frac{z^2}{2N^{1/2}} + \frac{z^3}{3N^{3/2}} \end{cases} \Rightarrow p \approx n k_B T \left(1 \mp \frac{n \lambda_T^3}{4N^{1/2}} + \left(\frac{1}{8} - \frac{2}{9N^{1/2}} \right) n^2 \lambda_T^3 \right)$$

* Working

$$\frac{x + \alpha x^2}{x + \beta x^2} \approx \frac{1 + \alpha x}{1 + \beta x}$$

$$\approx (1 + \alpha x)(1 - \beta x)$$

$$\hookrightarrow (1 \pm x/4N^{1/2})(1 \mp x/4N^{1/2})$$

$$\approx 1 \pm x/4N^{1/2} \mp x^2/16N$$

$$\approx 1 \pm \frac{x}{4N^{1/2}} \mp \frac{x^2}{16N}$$

Box-Einstein condensation

For (mu) bosons we considered

$\rightarrow z \leq 1$ since $\beta > 0, \mu \leq \epsilon_0 = 0$
 $\rightarrow$ else series diverges (for $z > 1$)

with $z = e^{\beta \mu}, n = N/V, \log \Xi = \sum_{r=1}^{\infty} \frac{1}{r} \frac{z^r}{1 - \exp(\beta(\mu - \epsilon_r))}$

$$n \lambda_T^3 = \sum_{l=1}^{\infty} \frac{z^l}{l^{3/2}} \quad \text{inverting}$$

$$\text{Now } \mu = \beta^{-1} \log z \rightarrow \mu(N, V, T)$$

Then $P(N, V, T)$ by replacing $z = e^{\beta \mu}$ by pressure since $\frac{PV}{k_B T} = \frac{V}{\lambda_T^3} \sum_{l=1}^{\infty} \frac{z^l}{l^{5/2}}$

$$\rightarrow n \lambda_T^3 = \sum_{l=1}^{\infty} \frac{z^l}{l^{3/2}} \quad \text{converges for } z \leq 1, \text{ diverges for } z > 1$$

$$\text{for } z \rightarrow 1, \mu \rightarrow \epsilon_0: n \lambda_T^3 = \sum_{l=1}^{\infty} \frac{1}{l^{3/2}} = \zeta(3/2) \approx 2.612 \dots$$

$\Rightarrow$ can only find a chem. potential for $n \lambda_T^3 \leq 2.612$
 else unphysical result (no inversion possible)
 (no μ)

$\rightarrow$ reevaluate analysis:

$$N = \sum_{\vec{p}} \langle n_{\vec{p}} \rangle = \sum_{\vec{p}} \frac{1}{e^{\beta(P^2/2m - \mu)} - 1} = \underbrace{\frac{z}{1-z}}_{\vec{p}=0} + \sum_{\vec{p} \neq 0} \frac{z}{e^{\beta P^2/2m} - z}$$

for $z \rightarrow 1$ the 1st term diverges

this isn't taken into account in the integral form (z.o.z.)

"operating ground state contrib from excited states"

$$n \lambda_T^3 = \frac{\lambda_T^3}{V} \frac{z}{1-z} + \sum_{l=1}^{\infty} \frac{z^l}{l^{3/2}}$$

* low n , high T ($n \lambda_T^3 < 2.612$):

"classical regime", can neglect 1st term $\sim O(1/V)$

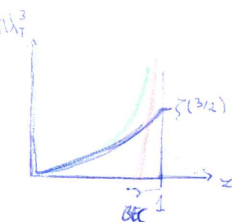
* high n , low T ($n \lambda_T^3 > 2.612$):

take into account ground state $\sim O(1)$

since a macroscopic number of particles can occupy the ground state

$\rightarrow$ BEC

correction term for ground state contributions not taken into account for fermions
 since only 1 fermion can occupy that state $\rightarrow$ minimal mistake



with $\zeta(s)$ Riemann zeta
 $\zeta(s) \equiv \sum_{l=1}^{\infty} \frac{1}{l^s}$
 $\zeta = 0$ for $s = -2n$ or $s = \frac{1}{2} + i\gamma$
 $n \in \mathbb{N}_0$ (Riemann Hypothesis)

$$N = \frac{z}{1-z} + \sum_{l=1}^{\infty} \frac{z^l}{l^{3/2}} \quad \text{or} \quad N = \frac{z}{1-z} + \frac{V}{\lambda_T^3} \sum_{l=1}^{\infty} \frac{z^l}{l^{3/2}}$$

estimate

$$N \lambda_r^3 \approx \zeta(3/2) \Rightarrow (N/V)^{2/3} \frac{h^2}{2\pi m k_B T} \approx (\zeta(3/2))^{2/3}$$

$$\Rightarrow k_B T_{BEC} \approx \frac{h^2}{2\pi m} \left(\frac{N}{V \zeta(3/2)} \right)^{2/3}$$

→ atoms/mol. superfluids at low T

general (also valid for non-interacting bosons subject to external pot.)
bosons w density of states $g(\epsilon)$

$$[g(\epsilon)d\epsilon = \# \text{ of states w energy } [\epsilon, \epsilon+d\epsilon]]$$

$$\rightarrow N = \sum_{\epsilon} \frac{1}{e^{\beta(\epsilon - \mu)} - 1}$$

$$\approx \int_{\epsilon_{\min}}^{+\infty} \frac{g(\epsilon)d\epsilon}{e^{\beta(\epsilon - \mu)} - 1}$$

convergent if $\mu < \epsilon_{\min}$ for $\mu \rightarrow \epsilon_{\min} \Rightarrow$ BECwhether this happens
~ $g(\epsilon)$ in vicinity $\epsilon_{\min}$

Photons and black body radiation

model: cubic cavity of side L

inside photons in thermal equilibrium at temp T

$$\rightarrow \text{Maxwell eq. } 1/c^2 \partial_t^2 \vec{E} = \nabla^2 \vec{E}$$

$$\text{solution: } \vec{E} = \vec{E}_0 \text{Re} \{ \exp(i(\vec{k} \cdot \vec{x} - \omega t)) \}$$

(traveling wave with speed c in direction $\vec{k}$)

$$\rightarrow \text{dispersion relation: } \omega = c|\vec{k}|$$

in a box $\Rightarrow$ boundary conditions: assume conducting walls

$$\vec{E} = \vec{E}_0 \sin(\vec{k} \cdot \vec{x}) \sin(\omega t)$$

$$\text{with } \vec{k} = \frac{\pi}{L} (n_x, n_y, n_z)$$

$$n_i \in \mathbb{N}_0, (x, y, z) \in (0, L)$$

then # of allowed waves in a volume element: $\frac{2V}{\pi^3} d\vec{k}$ photons have 2 transverse polarizations ($\vec{\nabla} \cdot \vec{E} = 0 \Rightarrow \vec{k} \cdot \vec{E} = 0$)

$$\Rightarrow \text{density of state: } g(\omega)d\omega \equiv \frac{1}{g} \frac{2V}{\pi^3} 4\pi k^2 dk = \frac{V}{\pi^2 c^3} \omega^2 d\omega$$

Planck: $\epsilon = \hbar\omega(n + 1/2)$

$$\rightarrow \Xi = \sum_{n, \vec{k}} \frac{1}{1 - e^{-\beta \hbar \omega}} = \prod_{\vec{k}} \frac{1}{1 - e^{-\beta \hbar \omega_{\vec{k}}}}$$

→ bosons with $\mu = 0$

$$\rightarrow \log \Xi = - \int_0^\infty d\omega g(\omega) \log(1 - e^{-\beta \hbar \omega})$$

$$E = - \frac{\partial \log \Xi}{\partial \beta} = \int_0^\infty d\omega g(\omega) \frac{\hbar \omega}{e^{\beta \hbar \omega} - 1}$$

x = $\beta \hbar \omega$ (dimensionless)

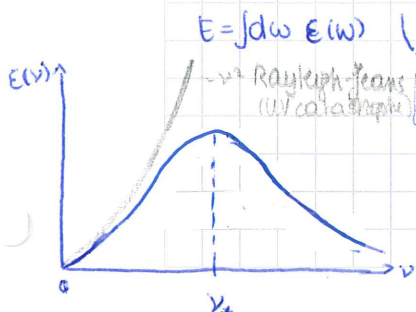
$$= \frac{V \hbar}{\pi^2 c^3} \int_0^\infty d\omega \frac{\omega^3}{e^x - 1} = \frac{V \hbar}{\pi^2 c^3} \frac{1}{\beta^4 \hbar^4} \int_0^\infty dx \frac{x^3}{e^x - 1}$$

$$= \frac{V \hbar}{c^3} \left(\frac{k_B T}{\hbar} \right)^4 \left(\frac{\pi}{\pi^2} \right) = \frac{V \pi^2 \hbar}{15 c^3} \left(\frac{k_B T}{\hbar} \right)^4$$

Stefan-Boltzmann: $E \sim T^4$ ($E \propto T^4$)

$$I = \int_0^\infty dx \frac{x^3}{e^x - 1} = \int_0^\infty dx x^3 \sum_{n=1}^\infty e^{-nx} = \sum_{n=1}^\infty \int_0^\infty dx x^3 e^{-nx} \rightarrow \text{3 times integ. by parts}$$

$$= \sum_{n=1}^\infty \int_0^\infty \frac{6 e^{-nx}}{n^3} dx = \sum_{n=1}^\infty \frac{6}{n^4} \left(-\frac{1}{n^4} \right) \Big|_0^\infty = \sum_{n=1}^\infty \frac{6}{n^4} = 6 \sum_{n=1}^\infty \frac{1}{n^4} = 6 \zeta(4) = \frac{\pi^4}{15}$$



$$E(\omega) = \frac{\hbar}{\pi^2 c^3} \frac{\omega^3}{e^{\beta \hbar \omega} - 1}$$

Planck's law

→ energy dens per unit volume

expand

$$\omega \gg 1 \rightarrow E \sim 0$$

$$\omega \ll 1 \rightarrow x \sim 0: \frac{x^3}{e^x - 1} \approx x^2 - \frac{x^3}{2}$$

$$\Rightarrow E_p \approx \frac{1}{\pi^2 c^3} \left\{ \omega^2 k_B T - \frac{1}{2} \hbar \omega^3 + \dots \right\}$$

equip. quantum con.

Fermions

non-interacting fermions:

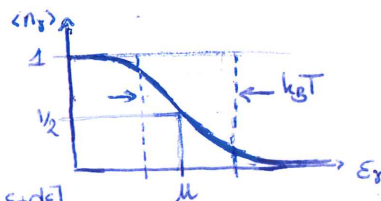
$$\log \Xi = \sum_Y \log (1 + e^{-\beta(\epsilon_Y - \mu)})$$

$$\rightarrow E = -\frac{\partial \log \Xi}{\partial \beta} \Big|_{\beta \mu} = \sum_Y \frac{\epsilon_Y}{e^{\beta(\epsilon_Y - \mu)} + 1} = \sum_Y \langle n_Y \rangle \epsilon_Y \approx \int_0^\infty \frac{\epsilon q(\epsilon) d\epsilon}{1 + e^{\beta(\epsilon_Y - \mu)}}$$

$$\langle n_Y \rangle = \frac{1}{1 + e^{\beta(\epsilon_Y - \mu)}}$$

Fermi function

$q(\epsilon) = \text{dens. of states}$
 $q(\epsilon) d\epsilon = \# \text{ states in } [\epsilon, \epsilon + d\epsilon]$



fermions $\rightarrow \mu$ is not bounded, can take any possible value

* for $\epsilon_0 = 0$: if $\mu < 0$ and $-\beta\mu \gg 1 \Rightarrow e^{-\beta\mu} \gg 1$

$$\rightarrow \langle n_Y \rangle \approx e^{-\beta(\epsilon_Y - \mu)}$$

Fermi $\rightarrow$ classical

if $\mu > 0$ and $\beta\mu \gg 1 \Rightarrow$ distribution becomes sharper

* for $T = 0$: all states below Fermi energy ϵ_F occupied
 $\rightarrow$ degenerate Fermi system

$$E_0 = \int_0^{\epsilon_F} \epsilon q(\epsilon) d\epsilon \quad \text{and} \quad N = \int_0^{\epsilon_F} q(\epsilon) d\epsilon$$

define Fermi temperature $T_F = \epsilon_F / k_B$: $T \ll T_F \rightarrow \langle n_Y \rangle$ is very sharp and almost fully degenerate

for $T \rightarrow 0$

$$\langle n_Y \rangle = \begin{cases} 1 & \epsilon_Y < \epsilon_F \\ 1/2 & \epsilon_Y = \epsilon_F \\ 0 & \text{else} \end{cases} \quad \text{with } \epsilon_F \equiv \lim_{T \rightarrow 0} \mu$$

$$\text{thus } \lim_{T \rightarrow 0} E \equiv E_0 = \int_0^{\epsilon_F} \epsilon q(\epsilon) d\epsilon$$

$$\lim_{N \rightarrow 0} N = \int_0^{\epsilon_F} q(\epsilon) d\epsilon$$

$$\frac{PV}{k_B T} = \log \Xi \approx \int_0^\infty d\epsilon q(\epsilon) \log(1 + e^{-\beta(\epsilon_Y - \mu)})$$

$$\Leftrightarrow \lim_{T \rightarrow 0} p = \lim_{T \rightarrow 0} \frac{1}{V} \int_0^{\epsilon_F} d\epsilon q(\epsilon) (\epsilon_F - \epsilon)$$

$$= \frac{1}{V} \int_0^{\epsilon_F} d\epsilon q(\epsilon) (\epsilon_F - \epsilon) = p_0 > 0$$

$$\lim_{T \rightarrow 0} \frac{1}{\beta} \log(1 + e^{-\beta(\epsilon_Y - \mu)}) = \begin{cases} -(\epsilon_Y - \epsilon_F) & \epsilon < \epsilon_F \\ 0 & \epsilon > \epsilon_F \end{cases}$$

$$\text{for } e^{\text{large}}: \log(1 + e^{\beta \dots}) \approx \log(e^{\beta \dots}) = \beta \dots$$

$\rightarrow$ pressure does NOT vanish

some fermions have non-zero momentum due to the exclusion principle
 $\Rightarrow$ non-zero pressure

[in their ground state: occupy high mom. state $\rightarrow$ generate finite pressure]

low temperature behaviour

$$E = \int_0^\infty d\epsilon \frac{q(\epsilon) \epsilon}{e^{\beta(\epsilon - \mu)} + 1} = \int_0^\mu d\epsilon q(\epsilon) \frac{1 + e^{\beta(\epsilon - \mu)} - e^{\beta(\epsilon - \mu)}}{1 + e^{\beta(\epsilon - \mu)}} \epsilon + \int_\mu^\infty d\epsilon \frac{q(\epsilon) \epsilon}{1 + e^{\beta(\epsilon - \mu)}}$$

$$= \int_0^\mu d\epsilon q(\epsilon) \epsilon - \int_0^\mu d\epsilon \frac{q(\epsilon) \epsilon}{1 + e^{-\beta(\epsilon - \mu)}} + \int_\mu^\infty d\epsilon \frac{q(\epsilon) \epsilon}{1 + e^{\beta(\epsilon - \mu)}}$$

$$= \int_0^\mu d\epsilon q(\epsilon) \epsilon - \frac{1}{\beta} \int_0^\mu dx \frac{(\mu - x/\beta) q(\mu - x/\beta)}{1 + e^x} + \frac{1}{\beta} \int_0^\infty dx \frac{(\mu + x/\beta) q(\mu + x/\beta)}{1 + e^x}$$

$$= \int_0^\mu d\epsilon q(\epsilon) \epsilon + \frac{1}{\beta} \int_0^\infty dx \frac{(\mu + x/\beta) q(\mu + x/\beta) - (\mu - x/\beta) q(\mu - x/\beta)}{1 + e^x}$$

sub with
 $x \equiv \beta(\epsilon - \mu)$

at $q(\epsilon) = 0$ for $\epsilon < 0$
 assuming that $\epsilon_0 = 0$

$\rightarrow$ can extend 1st integr.
 without any problems

$$E = \int_0^{\mu} d\epsilon q(\epsilon) \epsilon + k_B T \int_0^{\infty} dx \frac{(\mu + k_B T x) q(\mu + k_B T x) - (\mu - k_B T x) q(\mu - k_B T x)}{1 + e^x}$$

low temp T: expansion

$$q(\mu \pm k_B T x) \approx q(\mu) \pm k_B T x q'(\mu)$$

$$\Rightarrow (\mu \pm k_B T x) q(\mu \pm k_B T x) \approx \mu q(\mu) \pm k_B T x (\mu q'(\mu) + q(\mu))$$

(only keep smallest order term in $k_B T$)

$$\Rightarrow E \approx \int_0^{\mu} d\epsilon q(\epsilon) \epsilon + 2(k_B T)^2 (\mu q'(\mu) + q(\mu)) \int_0^{\infty} \frac{x dx}{1 + e^x}$$

as $T \rightarrow 0$

$\mu \rightarrow \epsilon_F$

$$= \int_0^{\epsilon_F} d\epsilon q(\epsilon) \epsilon + \int_{\epsilon_F}^{\mu} d\epsilon q(\epsilon) \epsilon + \frac{\pi^2}{6} (k_B T)^2 [\epsilon_F q'(\epsilon_F) + q(\epsilon_F)]$$

$$\int_0^{\infty} \frac{x dx}{1 + e^x} = \frac{\pi^2}{12}$$

$$\int_{x+\Delta x}^{x+\Delta y} f(x) dx \approx \int_{x+\Delta x}^{x+\Delta y} [f(x_0) + f'(x_0)(x-x_0) + \dots] dx$$

$$\approx \int_0^{\epsilon_F} d\epsilon q(\epsilon) \epsilon + q(\epsilon_F) \epsilon_F (\mu - \epsilon_F) + (k_B T)^2 [\epsilon_F q'(\epsilon_F) + q(\epsilon_F)] \frac{\pi^2}{6}$$

similarly

$$N \approx \int_0^{\mu} d\epsilon \frac{q(\epsilon)}{1 + e^{\beta(\epsilon - \mu)}} \approx \dots \approx \int_0^{\epsilon_F} d\epsilon q(\epsilon) + q(\epsilon_F) (\mu - \epsilon_F) + \frac{\pi^2}{6} (k_B T)^2 q'(\epsilon_F)$$

$$\Leftrightarrow (\mu - \epsilon_F) q(\epsilon_F) = -\frac{\pi^2}{6} q'(\epsilon_F) (k_B T)^2 + O((N - N_0))$$

$$\Rightarrow E \approx E_0 - \epsilon_F \frac{\pi^2}{6} q'(\epsilon_F) (k_B T)^2 + \frac{\pi^2}{6} (k_B T)^2 \epsilon_F q'(\epsilon_F) + \frac{\pi^2}{6} (k_B T)^2 q(\epsilon_F)$$

$$= E_0 + \frac{\pi^2}{6} (k_B T)^2 q(\epsilon_F)$$

$$\Rightarrow C_V = + \frac{\partial E}{\partial T} \Big|_V \approx \frac{\pi^2}{3} q(\epsilon_F) k_B^2 T$$

which vanishes for low T

in good approx: conducting e^- in metal $\sim$ gas of free Fermions

@ T_{room} : C_V contribution e^- masked by that of lattice vibrations ($\sim T^3$)

@ T_{low} : $C_V \sim T$ e^- , no longer effect lattice vibr. (phonons)

$q(\epsilon)$ for free particles:

$$\vec{p} = \hbar \vec{n} / L$$

$\Rightarrow$ #states with mom. in $[p, p+dp]$

$$2 \frac{V}{h^3} 4\pi p^2 dp = \frac{16\pi V}{h^3} m^{3/2} \sqrt{2\epsilon} d\epsilon \equiv q(\epsilon) d\epsilon$$

Oral exam questions

- 1] Show how the average number of particles $\langle N \rangle$ and the variance of N $\sigma_N = \langle N^2 \rangle - \langle N \rangle^2$ can be calculated from the partition function Ξ .

Now consider a system of non-interacting particles and assume that the one-particle $\chi_1(V, T)$ is known. Compute Ξ and show that the relative fluctuations of the number of particles is small.

- 2] Discuss the formula $n\lambda_T^3 = \frac{\lambda_T^3}{V} \frac{z}{1-z} + \sum_{l=1}^{\infty} \frac{z^l}{l^{3/2}}$ where $z = e^{\beta\mu}$ (Bosons) in the 2 limits

- with low density and high temp; $n\lambda_T^3 \gg 1$
- with high dens and low temp; $n\lambda_T^3 \ll 1$.

Pay in particular attention to the ground state and discuss some physical systems where this analysis is applicable and important

- 3] In the mean field approx we have the equation $m = \tanh(\beta(mJz + H))$. Explain the meaning of every quantity in the formula. Show that there is a critical temperature so that there ~~is~~ is spontaneous magn. for temp lower than T_c . In the vicinity of the critical point, many things behave as a power law. Show this for one or two examples.

- 4] The equipartition theorem states that $\langle x_i \frac{\partial H}{\partial x_i} \rangle = k_B T \delta_{i,0}$. Explain the meaning of this formula. Show the meaning of "Every quadratic degree of freedom contributes $k_B T/2$ to the system." Show some applications to a few examples.

- 5] For bosons and fermions $\log \Xi = \mp \sum_i \log(1 \mp e^{\beta(\mu - \epsilon_i)})$. Derive this eq. and explain the + and - sign. Why would you want to use the grand canonical ensemble? Then derive from the result the average occupation number $\langle n_i \rangle$. Why is the AIT important in your derivation and in your end result?

- 6] Prove the ideal gas law $PV = Nk_B T$ in the canonical and grand-canonical ensemble. In general the ensembles are equivalent to each other in the limit that relative fluctuations $\sigma^2 / \langle x \rangle^2$ vanish. Give an example of this for an ideal gas in the canonical and grand-canonical ensemble

- 7] Sketch the derivation of Planck's law: $\epsilon(\omega) = \frac{\hbar}{\pi^2 c^3} \frac{\omega^3}{e^{\beta \hbar \omega} - 1}$. What does this mean? The density of states is the following $g(\omega) = \frac{V}{\pi^2 c^3} \omega^2$. How is your unit dependent on the temp? How would you derive $\epsilon(\omega)$ classically?

- 8] The pressure of a quantum gas of non-interacting particles is given by $P = nk_B T \left(1 \pm \frac{n\lambda_T^3}{4\sqrt{\pi}} \right)$. Why are there 2 different signs and what do they correspond to? The formula describes the effect of quantum corrections to the ideal gas law. These corrections are small if $n\lambda_T^3$ is small. Why? Which kind of dimensionless param. do you expect to control the quantum corrections for a gas in 2D?

- 9] The virial expansion for an interacting system reads $P = nk_B T (1 + b_2 n + b_3 n^2 + \dots)$ where b_i are the virial coefficients. The leading partner in b_2 usually depends on temp and it can be either pos or neg. Discuss the temp. dep. and the sign of b_2 . Do you know a system in which b_2 does not dep. on temp?

Written part

1) Consider particles in a potential $V(x) = -ax$ with diffusion coefficient D .

Show that $c(x,t) = \frac{N}{\sqrt{4Dt}} e^{-1(x-vt)^2/4Dt}$ is a solution of the drift-diffusion eq. and find v . Plot this solution for different times. Show that $F = \gamma v$ where F is the force on the particles and γ is the friction coeff.

2) Consider a single 1D oscillator in the canonical ensemble and calculate the average energy $\langle E \rangle$ and the variance of E , $\sigma_E^2 = \langle E^2 \rangle - \langle E \rangle^2$. Do the same for N indep. HO.

3) Consider a 2D oscillator in polar coord $\mathcal{H} = \frac{1}{2m} (p_r^2 + \frac{p_\theta^2}{r^2}) + \frac{m\omega^2 r^2}{2}$

1. Calculate the part. function in " "
2. Calculate $\langle E \rangle$ and compare with the result in Cartesian coord
3. Do this calculation with the equip. theorem in polar coords.

4) We have N part in a volume V at temp T . Consider a part V_1 with N_1

1. Calculate $\langle N_1 \rangle$ and $\langle N_1^2 \rangle$
2. Calculate the variance and show that relative fluct. are small.

5) Consider $N-1$ coupled osc. such that $\mathcal{H} = \sum_{i=1}^N \frac{p_i^2}{2m} + \sum_{i=1}^{N-1} \frac{K}{2} (x_{i+1} - x_i)^2$

1. Calculate $\langle E \rangle$ and $\langle E^2 \rangle - \langle E \rangle^2$ and show that the relative fluct are small
2. Repeat for anharmonic osc $\mathcal{H} = \sum_{i=1}^N \frac{p_i^2}{2m} + \sum_{i=1}^{N-1} \frac{K}{2} (x_{i+1} - x_i)^4$

6) Consider a syst of N molecules in vol V and temp T . The molecules each consist of 3 atoms m_1, m_2, m_3 . The atoms interact through a potential

$$\psi(\vec{x}_1, \vec{x}_2, \vec{x}_3) = \frac{K}{2} (\vec{x}_1 - \vec{x}_2)^2 + \frac{K}{2} (\vec{x}_2 - \vec{x}_3)^2$$

1. Calculate Z . Also calculate the internal energy and C_V
2. Calculate $\langle (\vec{x}_1 - \vec{x}_2)^2 \rangle$, $\langle (\vec{x}_2 - \vec{x}_3)^2 \rangle$, $\langle (\vec{x}_3 - \vec{x}_1)^2 \rangle$
3. Calculate the pressure on the syst